

22. Seminar Radijske Komunikacije

Šum v radijskih komunikacijah

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Seznam prosojnic predavanja: Šum v radijskih komunikacijah

- 1 - Nesporazum slavnih znanstvenikov
- 2 - Spektralna gostota šuma
- 3 - Toplotno sevanje črnega telesa
- 4 - Sprejeta moč toplotnega šuma
- 5 - Toplotno ravnovesje
- 6 - Naravni izvori šuma
- 7 - Naravni šum neba
- 8 - Zgled šuma Sonca
- 9 - Razmerje signal/šum sprejemnika
- 10 - Šumna temperatura verige
- 11 - Šumno število ojačevalnika
- 12 - Povezava $F \leftrightarrow T$
- 13 - Šum slabilca
- 14 - Razmerje G/T
- 15 - Šum aktivnih gradnikov
- 16 - Šumni parametri tranzistorja
- 17 - Občutljivost GSM telefona
- 18 - Spremembi F in S/N
- 19 - Meritev občutljivosti sprejemnika
- 20 - Postopek vroče/hladno
- 21 - Merilnik šumnega števila
- 22 - Izračun pogostnosti napak BER
- 23 - BER različnih modulacij
- 24 - Vnaprejšnje popravljanje napak
- 25 - Fazni šum oscilatorja
- 26 - Leesonova enačba
- 27 - Šum $1/f$
- 28 - Kvaliteta rezonatorja
- 29 - Fazno-sklenjena zanka PLL
- 30 - Posledice faznega šuma
- 31 - Fazni šum brez približkov
- 32 - Širina Lorentzove spektralne črte
- 33 - Šum kot merilni signal
- 34 - Vir kriptografskega ključa
- 35 - Šumna kriptografija
- 36 - Psevdonaključna zaporedja LFSR
- 37 - Uporaba psevdonaključnih zaporedij

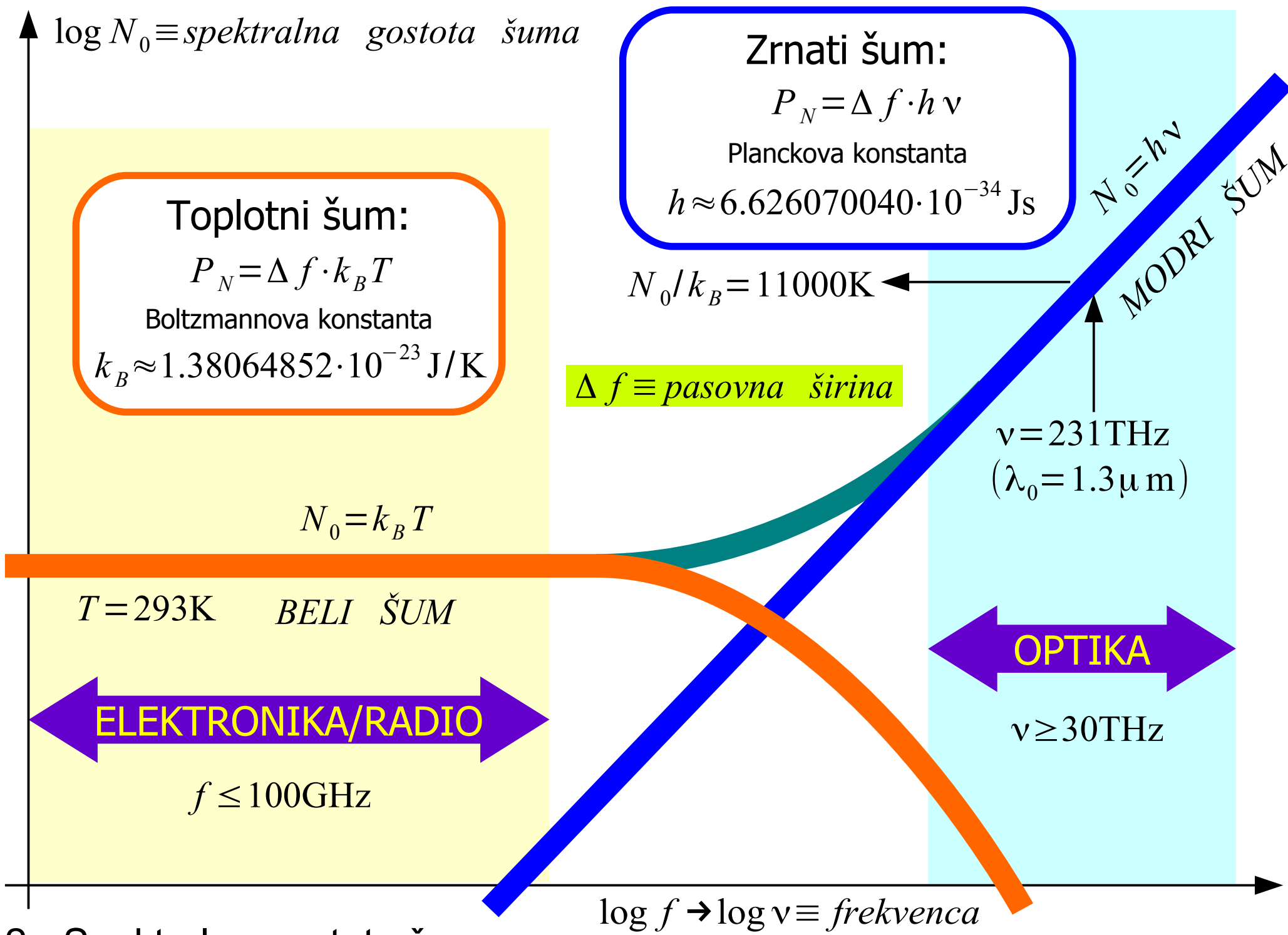
Peta mednarodna konferenca Solvay uglednih fizikov in kemikov na temo „Elektroni in fotoni“ (oktober 1927). Najuglednejša udeleženca Albert Einstein in Niels Bohr se nista razumela:

Albert Einstein: „God does not play dice!“
(Bog ne kocka! Torej ne dopušča naključnosti v naravnih zakonih.)

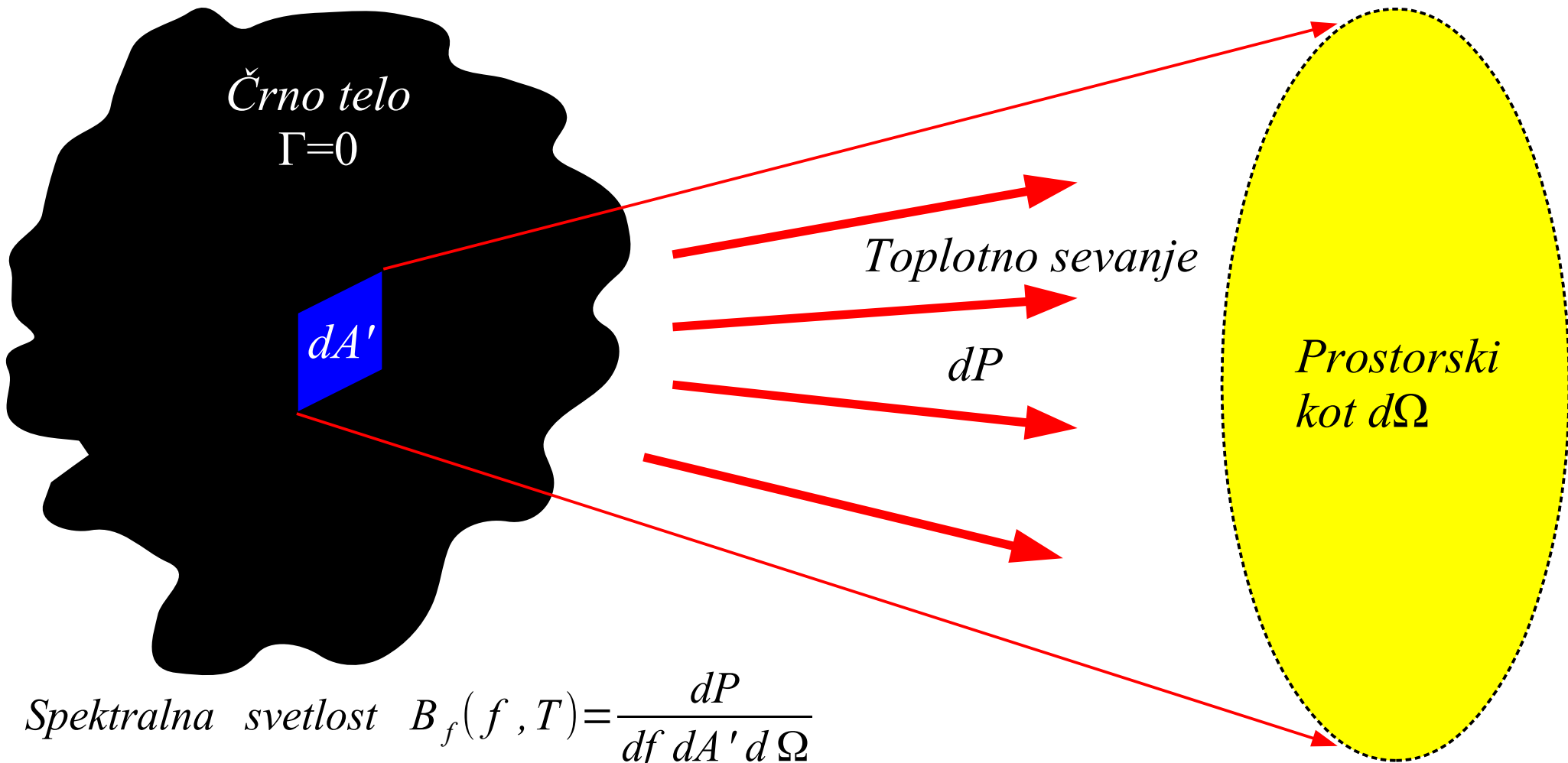
Niels Bohr: „Einstein, stop telling God what to do!“ (Einstein, nehaj učiti Boga, kaj naj počne!)

V telekomunikacijah naključnost imenujemo šum. Šum omejuje domet vsake zveze.

Šum je makroskopski opis kvantnih pojavov!



2 - Spektralna gostota šuma



Spektralna svetlost $B_f(f, T) = \frac{dP}{df dA' d\Omega}$

Planckov zakon $B_f(f, T) = \frac{2 h f^3}{c_0^2} \cdot \frac{1}{e^{\frac{h f}{k_B T}} - 1}$

Prazen prostor ϵ_0, μ_0
 $c_0 = 299792458 \text{ m/s} \approx 3 \cdot 10^8 \text{ m/s}$

Radio $h f \ll k_B T \rightarrow$ Rayleigh–Jeansov približek $B_f(f, T) \approx \frac{2 k_B T f^2}{c_0^2} = \frac{2 k_B T}{\lambda^2}$

3 – Toplotno sevanje črnega telesa

Prazen prostor ϵ_0, μ_0

Črno telo
 $\Gamma=0$

$$B_f = \frac{2k_B T}{\lambda^2}$$

dA'

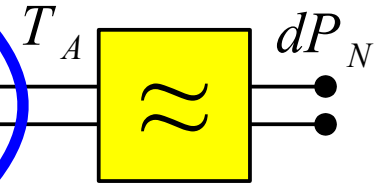
Samo ena polarizacija

Brezizgubna antena

$\eta=1$

$A_{eff}(\Theta, \Phi)$

$$dP_N = \frac{1}{2} \cdot B_f \cdot \Delta f \cdot dA' \cdot \Delta \Omega$$



Pasovno sito Δf

r

$$\Delta \Omega = \frac{A_{eff}(\Theta, \Phi)}{r^2} = \frac{\lambda^2 D(\Theta, \Phi)}{4\pi r^2} = \frac{\lambda^2 |F(\Theta, \Phi)|^2}{r^2 \iint_{4\pi} |F(\Theta^*, \Phi^*)|^2 d\Omega^*}$$

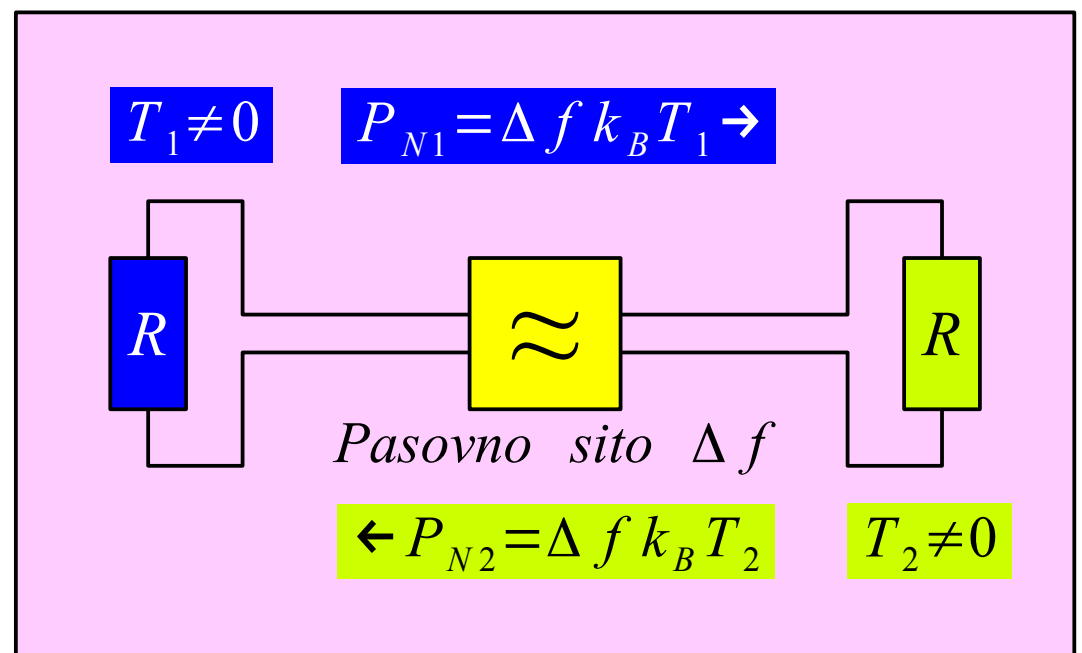
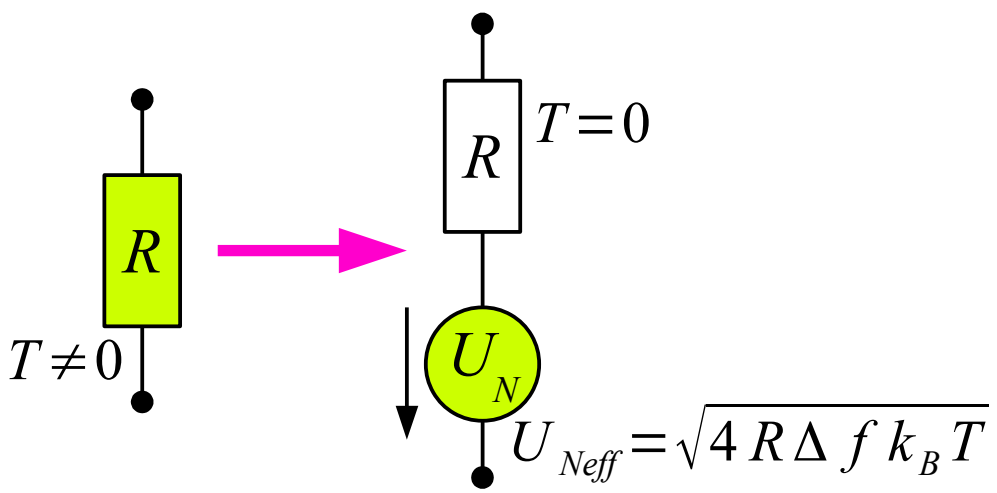
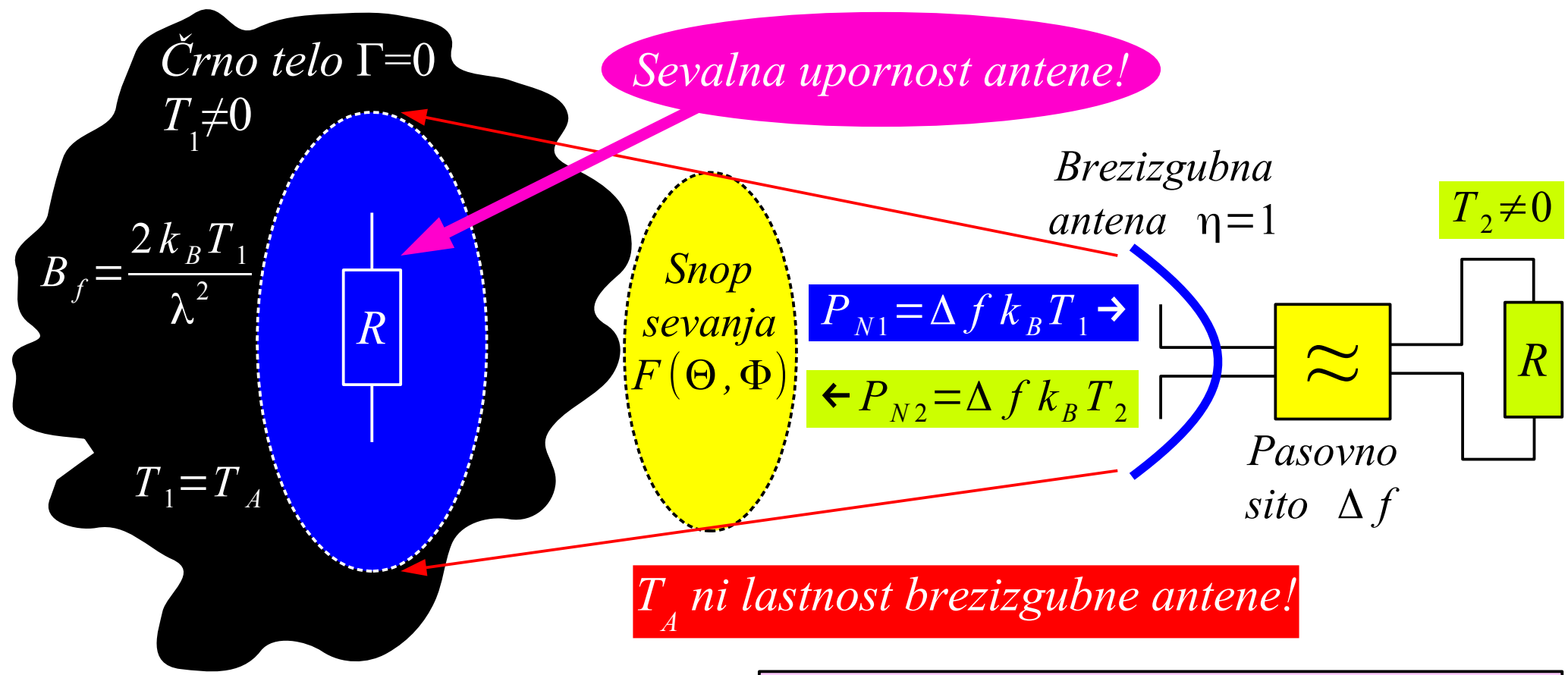
$$dA' = r^2 d\Omega$$

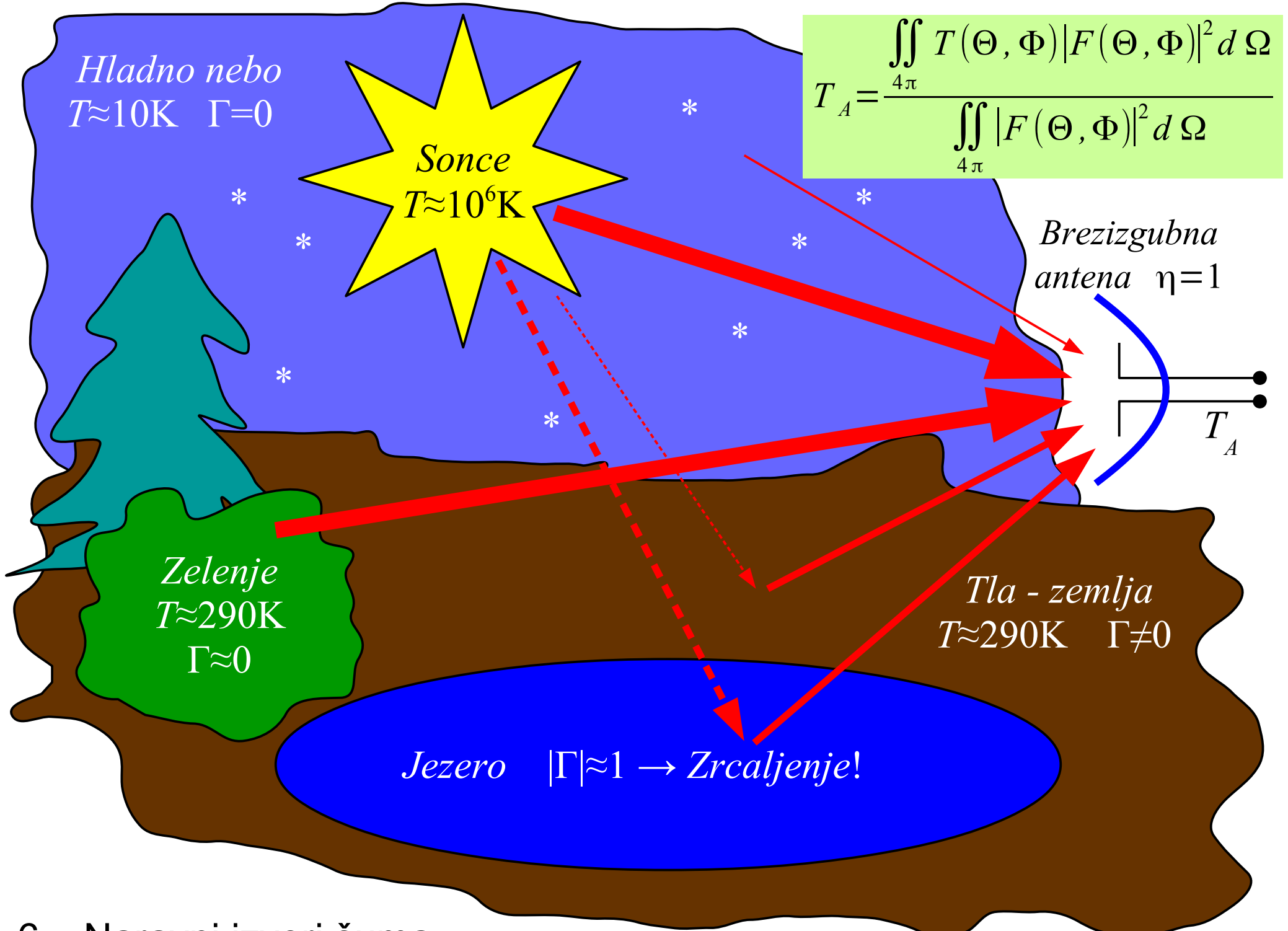
$$P_N = \iint_{A'} \frac{1}{2} \cdot B_f \cdot \Delta f \cdot dA' \cdot \Delta \Omega = \iint_{4\pi} \frac{1}{2} \cdot \frac{2k_B T(\Theta, \Phi)}{\lambda^2} \cdot \Delta f \cdot r^2 d\Omega \cdot \frac{\lambda^2 |F(\Theta, \Phi)|^2}{r^2 \iint_{4\pi} |F(\Theta^*, \Phi^*)|^2 d\Omega^*}$$

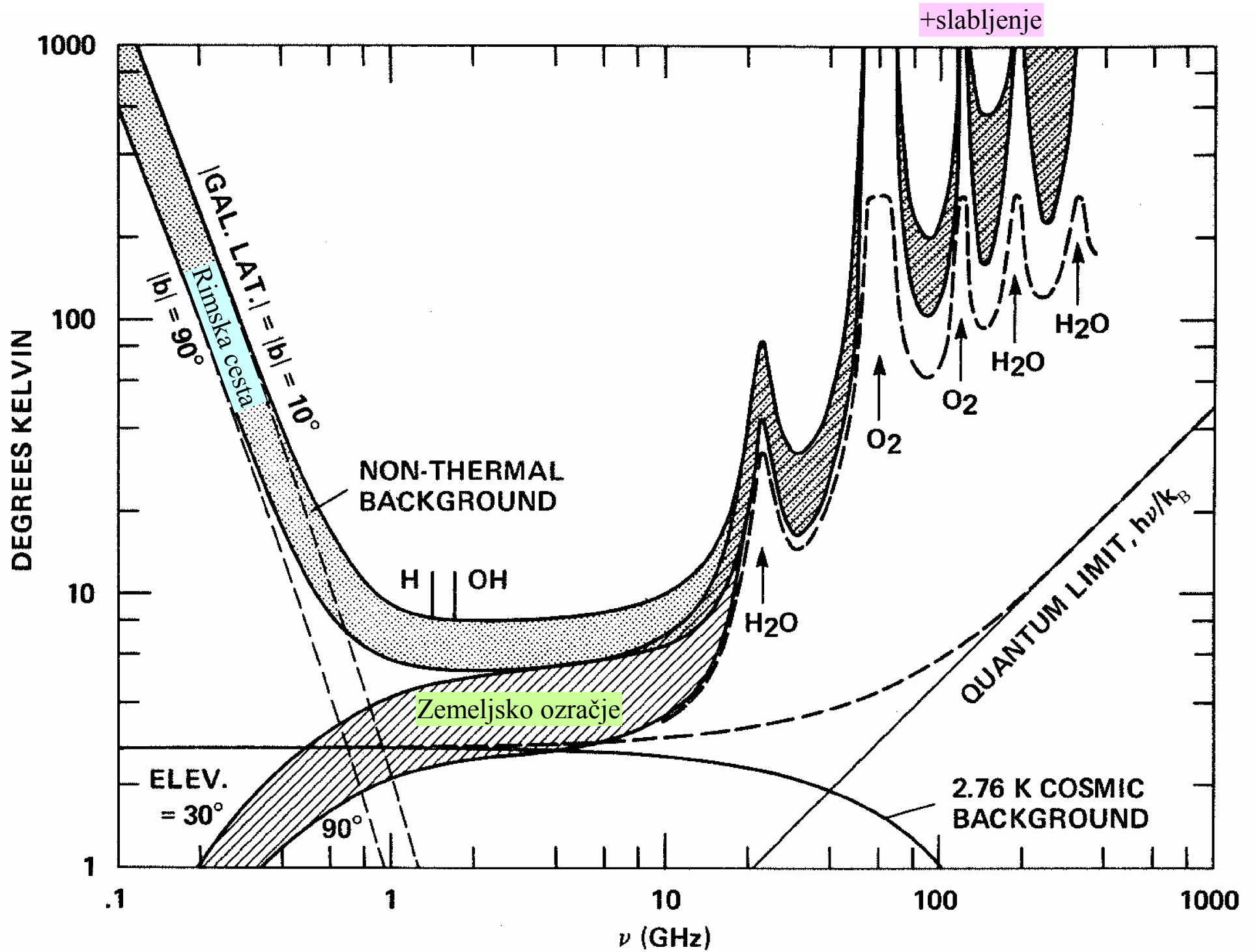
$$P_N = \Delta f k_B \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega} = \Delta f k_B T_A$$

$$T_A = \frac{\iint_{4\pi} T(\Theta, \Phi) |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

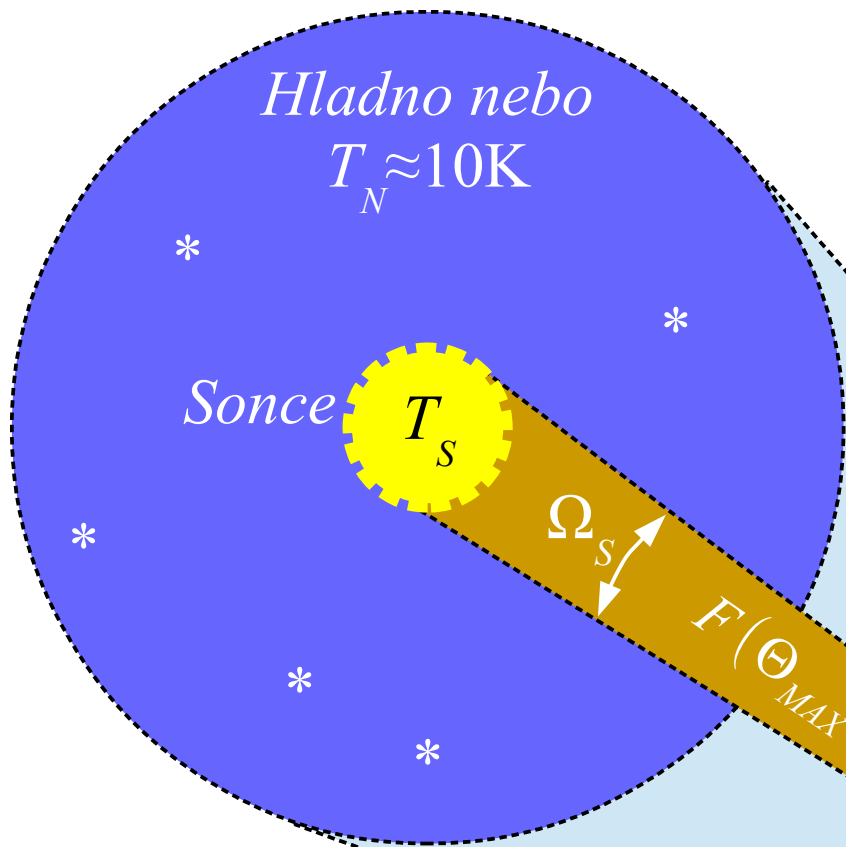
4 – Sprejeta moč toplotnega šuma







7 – Naravni šum neba



$$T_A = \frac{T_S \iint_{\Omega_S} |F(\Theta, \Phi)|^2 d\Omega + T_N \iint_{4\pi - \Omega_S} |F(\Theta, \Phi)|^2 d\Omega}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega}$$

$$D = \frac{4\pi |F(\Theta_{MAX}, \Phi_{MAX})|^2}{\iint_{4\pi} |F(\Theta, \Phi)|^2 d\Omega} \quad T_A \approx \frac{T_S \Omega_S D}{4\pi} + T_N$$

$$T_A \approx \frac{10^6 \text{ K} \cdot 6 \cdot 10^{-5} \text{ srd} \cdot 100}{4\pi \text{ srd}} + 10 \text{ K}$$

$$T_A \approx 476 \text{ K} + 10 \text{ K} = 486 \text{ K}$$

$$\alpha_S \approx 0.5^\circ \approx 9 \text{ mrd}$$

$$T_S \approx 10^6 \text{ K} @ f = 2 \text{ GHz}$$

$$\Omega_S = 2\pi [1 - \cos(\alpha_S/2)] \approx \pi \alpha_S^2 / 4 \approx 6 \cdot 10^{-5} \text{ srd}$$

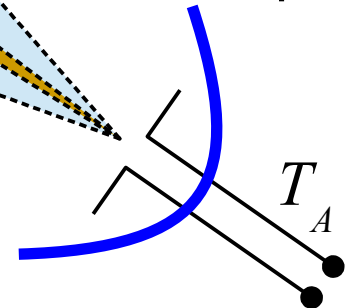
Zgled:
 $D = 20 \text{ dBi} = 100$

$$\Omega_A \approx \frac{4\pi}{D} = 0.126 \text{ srd} \gg \Omega_S$$

$F(\Theta, \Phi) \neq 0$

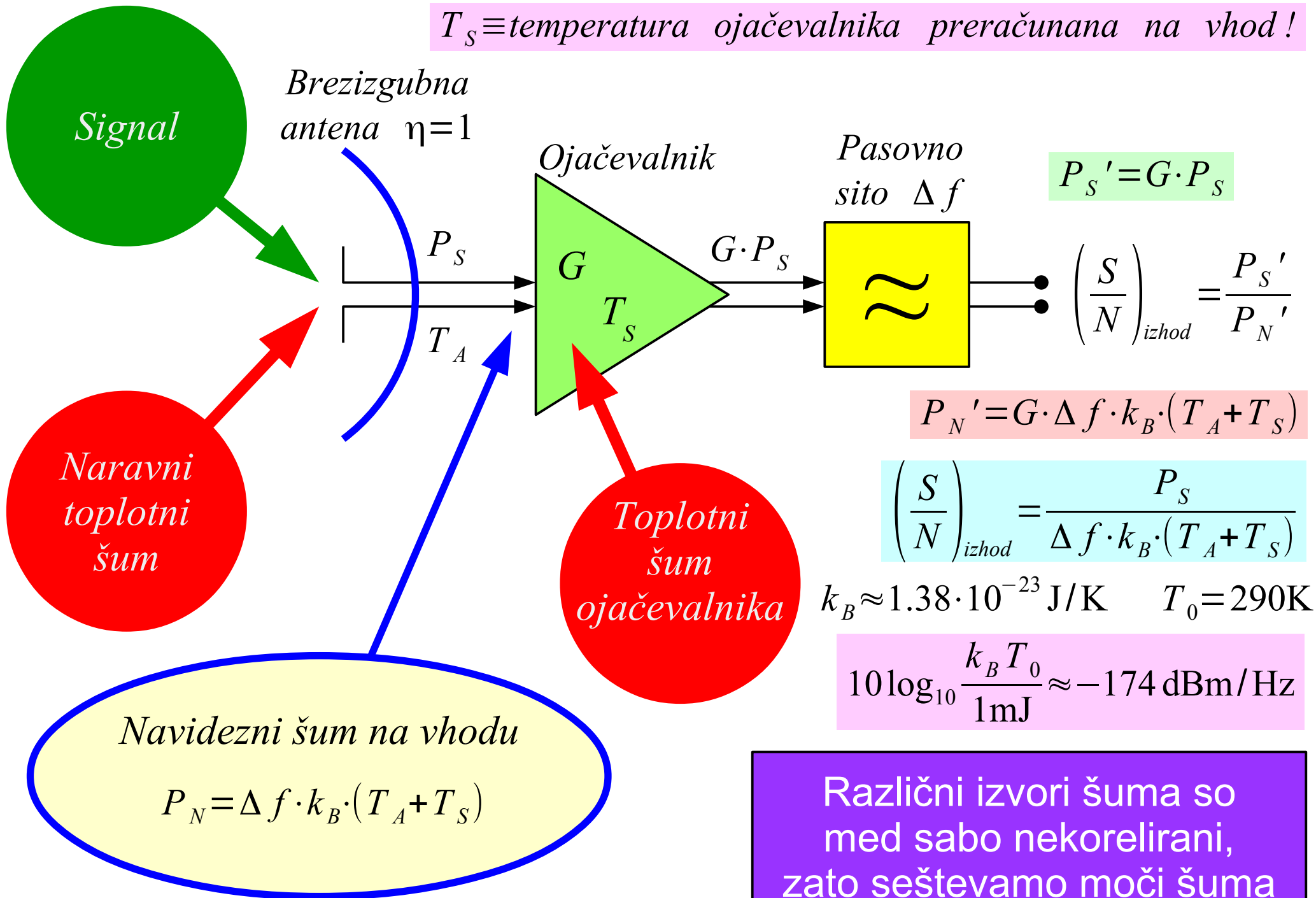
$F(\Theta, \Phi) = 0$
 izven snopa

Brezizgubna
 antenna $\eta = 1$



8 – Zgled šuma Sonca

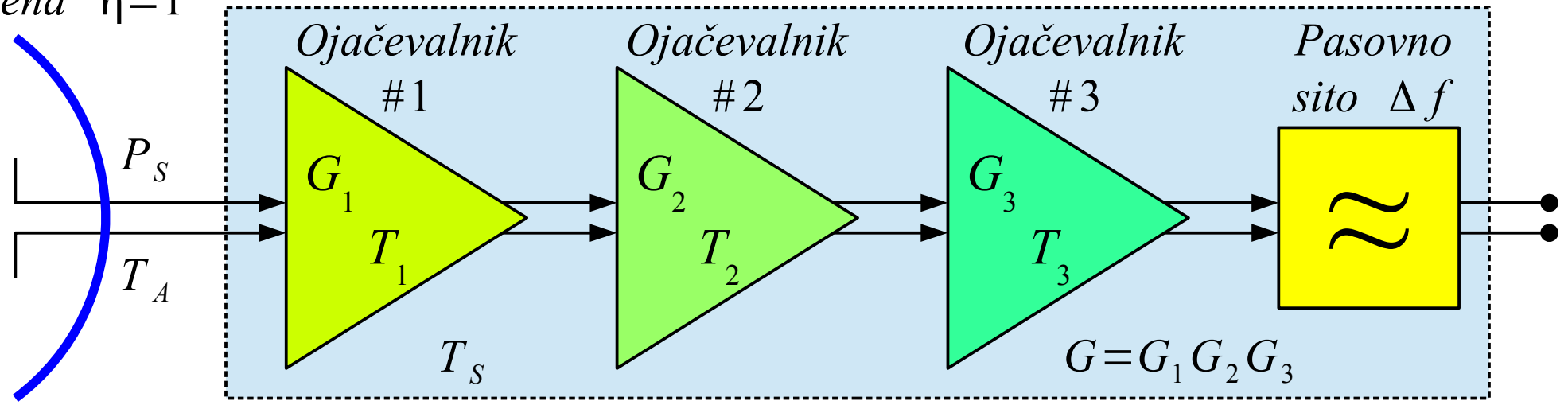
$T_S \equiv$ temperatura ojačevalnika preračunana na vhod!



Različni izvori šuma so med sabo nekorelirani, zato seštevamo moči šuma oziroma temperature šuma!

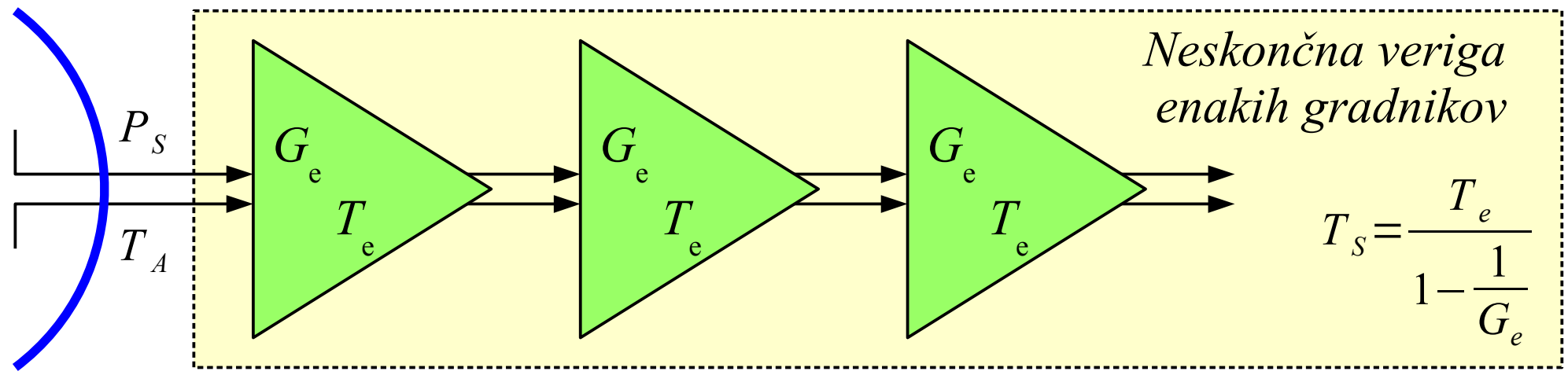
$$P_S' = G_3 G_2 G_1 P_S$$

Brezizgubna
antena $\eta=1$



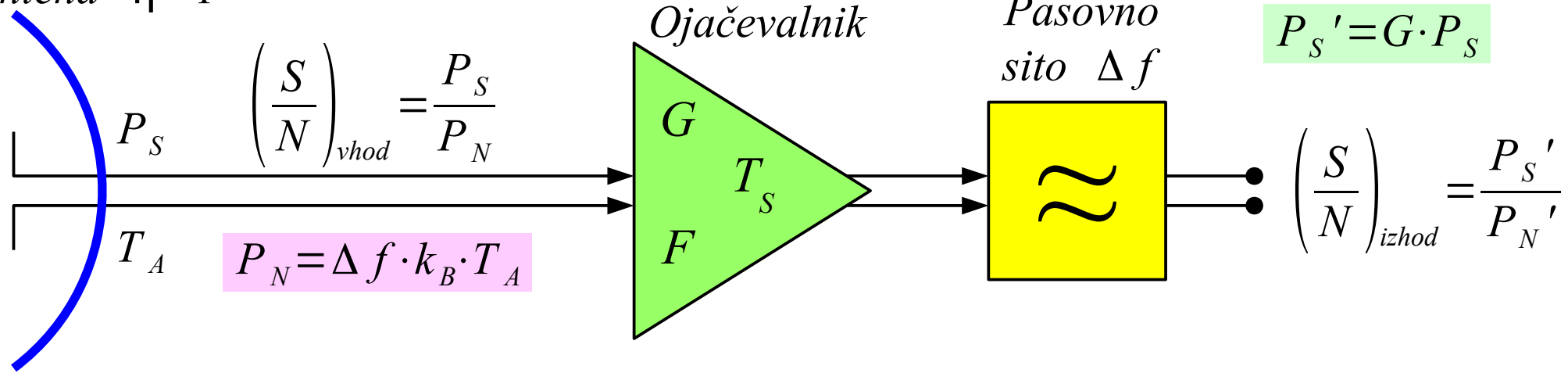
$$P_N' = \Delta f k_B [G_3 G_2 G_1 (T_A + T_1) + G_3 G_2 T_2 + G_3 T_3]$$

$$P_N' = G_3 G_2 G_1 \Delta f k_B (T_A + T_S) \quad \rightarrow \quad T_S = T_1 + \frac{T_2}{G_1} + \frac{T_3}{G_1 G_2} + \dots$$



$$T_S = \frac{T_e}{1 - \frac{1}{G_e}}$$

Brezizgubna
antena $\eta=1$



$$P_N' = G \cdot \Delta f \cdot k_B \cdot (T_A + T_S)$$

$$P_S' = G \cdot P_S$$

Nesmiselna definicija šumnega števila:

$$F = \frac{\left(\frac{S}{N}\right)_{\text{vhod}}}{\left(\frac{S}{N}\right)_{\text{izhod}}} = \frac{\frac{P_S}{\Delta f k_B T_A}}{\frac{G P_S}{G \Delta f k_B (T_A + T_S)}} = \frac{T_A + T_S}{T_A} = 1 + \frac{T_S}{T_A}$$

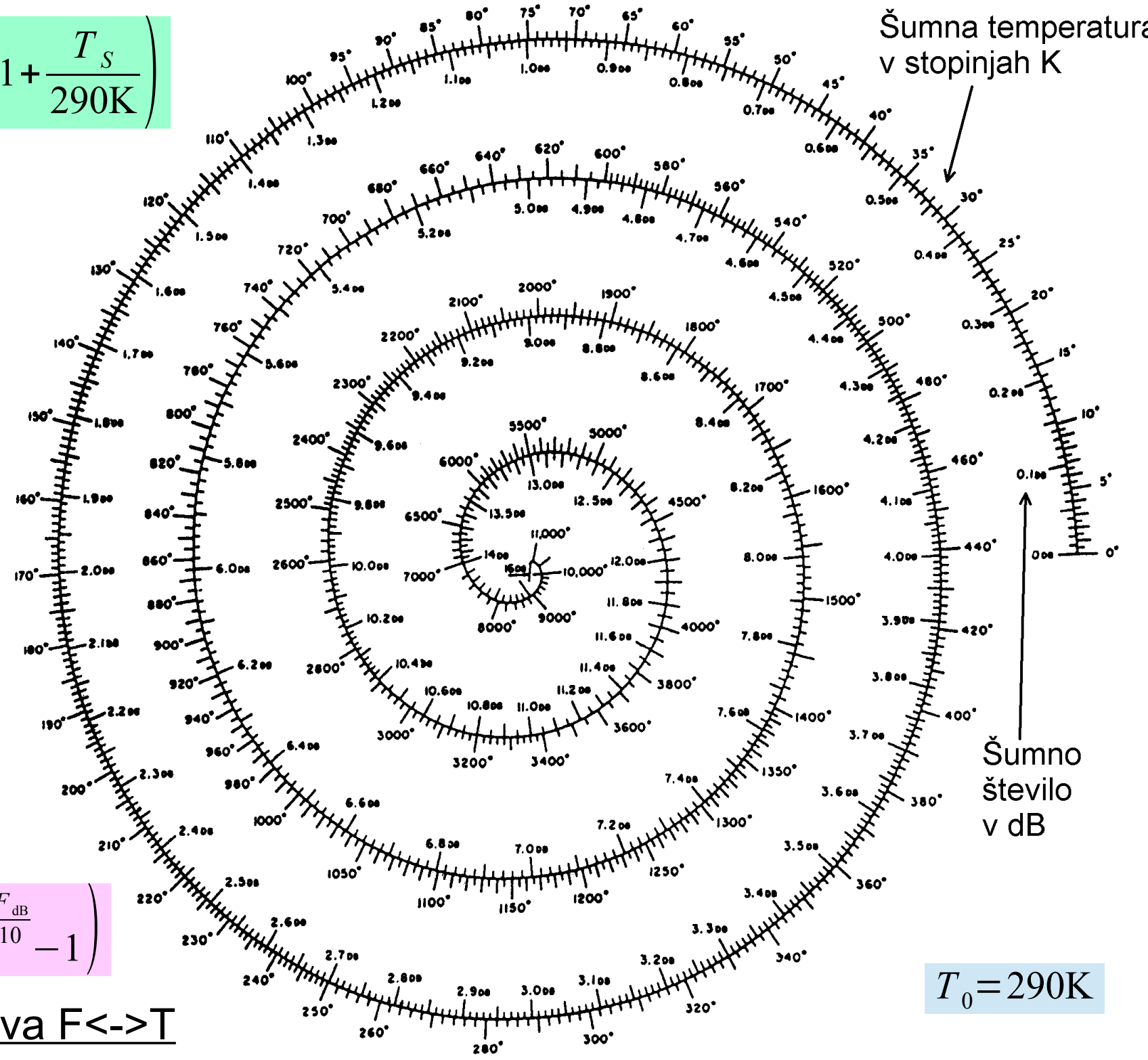
Lastnost ojačevalnika ne more biti funkcija T_A !

Smiselna definicija $F = 1 + \frac{T_S}{T_0} \quad @ \quad T_0 = 290\text{K} \quad \leftrightarrow \quad T_S = T_0 (F - 1)$

Logaritemske enote $F_{\text{dB}} = 10 \log_{10} F = 10 \log_{10} \left(1 + \frac{T_S}{T_0} \right) \quad \leftrightarrow \quad T_S = T_0 \left(10^{\frac{F_{\text{dB}}}{10}} - 1 \right)$

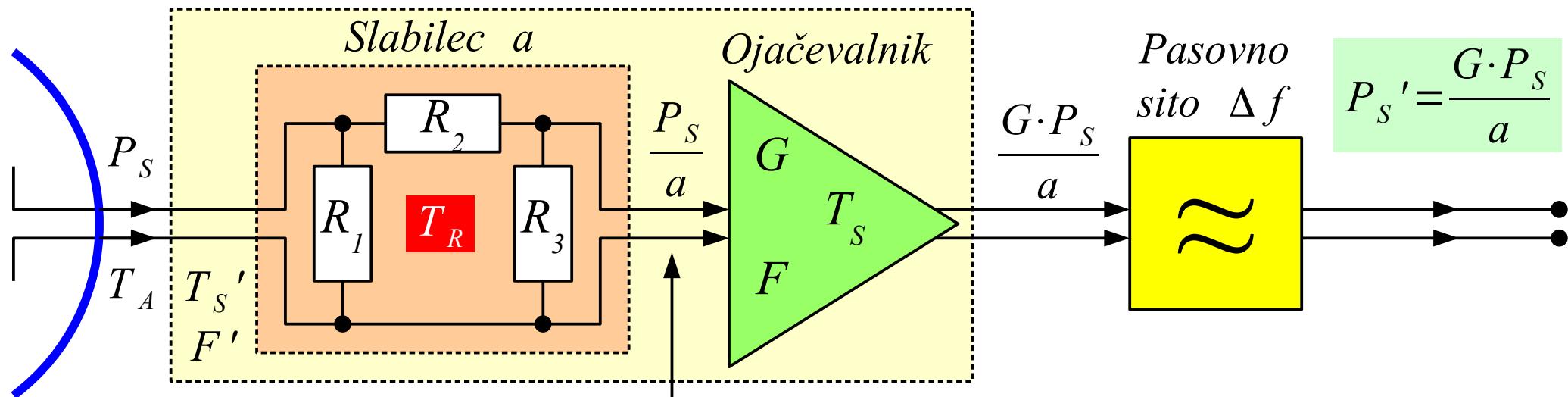
11 – Šumno število ojačevalnika

$$F_{dB} = 10 \log_{10} \left(1 + \frac{T_S}{290K} \right)$$



$$T_S = 290K \left(10^{\frac{F_{dB}}{10}} - 1 \right)$$

12 – Povezava F<->T



Brezizgubna antena $\eta=1$

$$\frac{T_A}{a} + T_R \left(1 - \frac{1}{a}\right)$$

$$P_N' = G \cdot \Delta f \cdot k_B \cdot \left[\frac{T_A}{a} + T_R \left(1 - \frac{1}{a}\right) + T_S \right]$$

$$T_S' = T_R(a-1) + a T_S$$

$$\left(\frac{S}{N}\right)_{\text{izhod}} = \frac{P_S'}{P_N'} = \frac{P_S}{\Delta f \cdot k_B \cdot [T_A + T_R(a-1) + a T_S]}$$

$$F' = 1 + \frac{T_S'}{T_0} = 1 + \frac{T_R}{T_0}(a-1) + a \frac{T_S}{T_0}$$

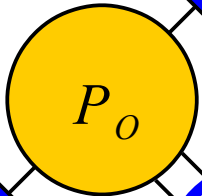
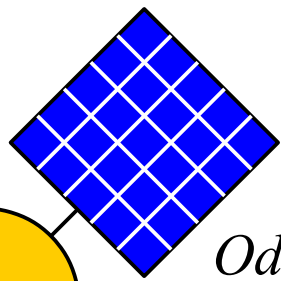
Pogost primer $T_R \approx T_0 = 290\text{K}$

$$F' \approx a + a \frac{T_S}{T_0} = a \left(1 + \frac{T_S}{T_0}\right) = a \cdot F$$

$$F_{\text{dB}}' \approx a_{\text{dB}} + F_{\text{dB}}$$

- Primeri slabilcev $T_R \approx T_0 = 290\text{K}$
 $F' \approx a \cdot F$ oziroma $F_{\text{dB}}' \approx a_{\text{dB}} + F_{\text{dB}}$
- (1) izgubna antena $a_{\text{dB}} = -10 \log_{10} \eta$
 - (2) prenosni vod z izgubami a_{dB}
 - (3) pasovno sito s slabljenjem a_{dB}
 - (4) slabljenje pasivnega mešalnika a_{dB}

Oddajnik na satelitu



Oddajna antena G_0

Zveza v praznem prostoru $P_S = P_0 \cdot G_0 \cdot G_S \cdot \left(\frac{\lambda}{4\pi r}\right)^2$

Oddajnik

Sprejemnik

$$\left(\frac{S}{N}\right)_{\text{izhod}} = P_0 \cdot G_0 \cdot \frac{1}{\Delta f \cdot k_B} \cdot \left(\frac{\lambda}{4\pi r}\right)^2 \cdot \frac{G_S}{(T_A + T_S)}$$

Sistem

Sprejemna postaja

$$(G/T) = \frac{G_S}{(T_A + T_S)} \text{ [K}^{-1}\text{]}$$

$$(G/T)_{\text{dB/K}} = 10 \log_{10} \frac{G_S \cdot 1\text{K}}{(T_A + T_S)} \text{ [dB/K]}$$

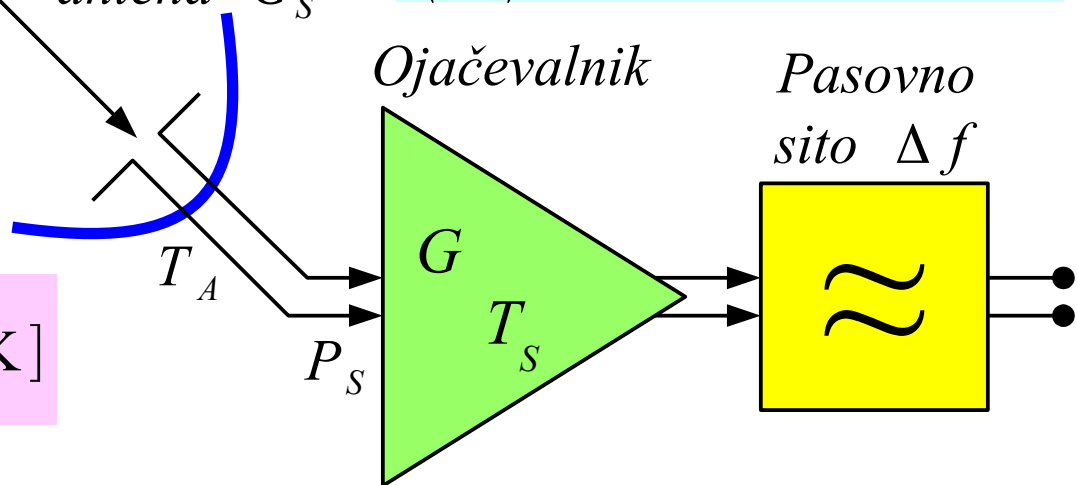
$$(G/T)_{\text{dB/K}} = G_{S\text{dB}} - 10 \log_{10} \frac{T_A + T_S}{1\text{K}} \text{ [dB/K]}$$

Sprejemna antena G_S

$$\left(\frac{S}{N}\right)_{\text{izhod}} = \frac{P_S}{\Delta f \cdot k_B \cdot (T_A + T_S)}$$

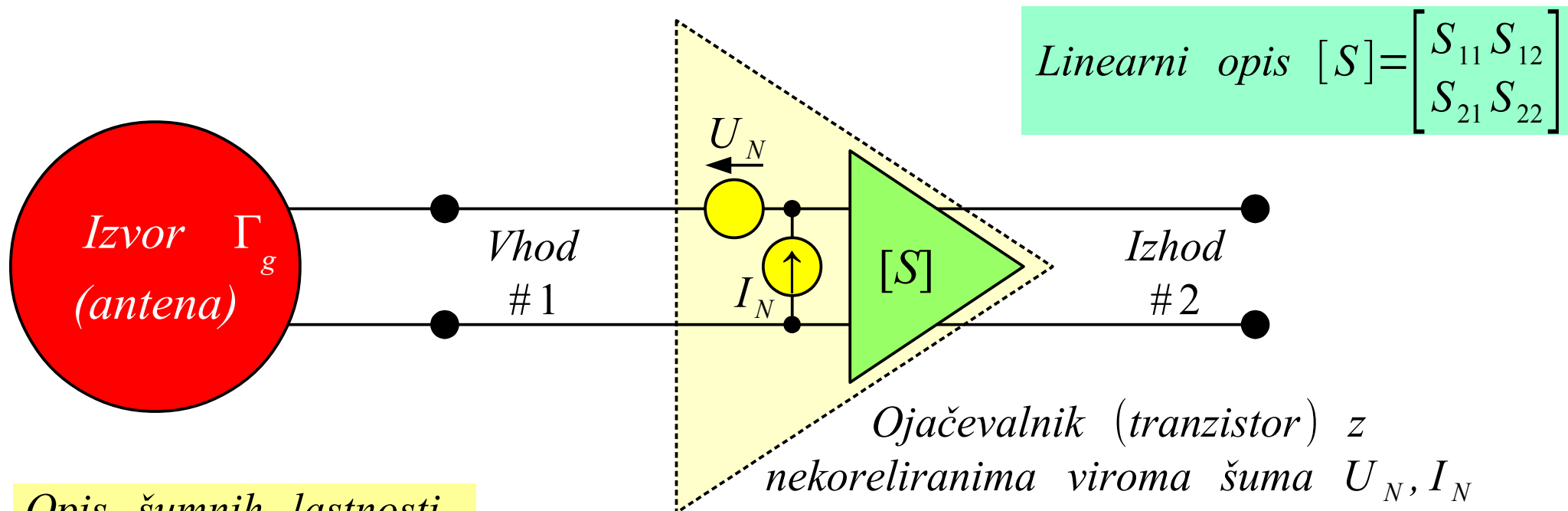
Ojačevalnik

Pasovno sito Δf



Zemeljska sprejemna postaja

Vrsta ojačevalnika	Ojačanje G [dB]	Temperatura šuma T_s [K]	Šumno število F_{dB} [dB]
Vakuumska cev z mrežicami (trioda, pentoda)	10↔20	1600↔9000	8↔15
Vakuumska cev s hitrostno modulacijo (klistron, TWT)	20↔50	3000↔30000	10↔20
Parametrični ojačevalnik (sobna temperatura)	10↔15	75↔300	1↔3
Si BJT, JFET ali MOSFET (sobna temperatura)	10↔20	75↔300	1↔3
GaAs FET ali HEMT (sobna temperatura)	10↔15	20↔120	0.3↔1.5
GaAs FET ali HEMT (hlajen 77K tekoči dušik)	10↔15	7↔35	0.1↔0.5
Si ali GaAs MMIC ojačevalnik	10↔25	170↔1600	2↔8
Operacijski ojačevalnik	40↔100	10^4 ↔ 10^9	16↔66



Opis šumnih lastnosti

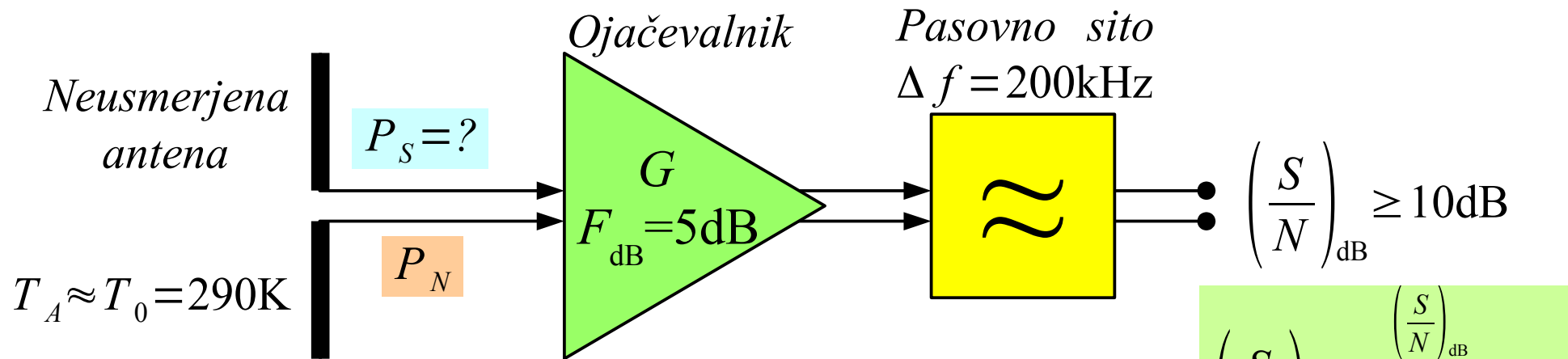
$U_N, I_N \rightarrow F_{MIN}, \Gamma_O, r_N$

$$F = F_{MIN} + 4 \frac{R_N}{Z_K} \cdot \frac{|\Gamma_g - \Gamma_O|^2}{(1 - |\Gamma_g|^2) \cdot |1 + \Gamma_O|^2} = F_{MIN} + 4 r_N \cdot \frac{|\Gamma_g - \Gamma_O|^2}{(1 - |\Gamma_g|^2) \cdot |1 + \Gamma_O|^2}$$

$F_{MIN} \equiv$ najnižje šumno število pri $\Gamma_g = \Gamma_O$ v linearnih enotah (ne v dB!)

$\Gamma_O \equiv$ optimalna odbojnost izvora za F_{MIN} (nima povezave z matriko $[S]$!)

$r_N = \frac{R_N}{Z_K} \equiv$ normirana šumna upornost (običajno $Z_K = 50 \Omega$)



$$T_S = T_0 \cdot \left(10^{\frac{F_{\text{dB}}}{10}} - 1\right) = 290\text{K} \cdot (3.162 - 1) = 627\text{K}$$

$$k_B \approx 1.38 \cdot 10^{-23} \text{ J/K}$$

$$\left(\frac{S}{N}\right) = 10^{\frac{\left(\frac{S}{N}\right)_{\text{dB}}}{10}} \geq 10$$

$$P_N = \Delta f \cdot k_B \cdot (T_A + T_S) = 200\text{kHz} \cdot 1.38 \cdot 10^{-23} \text{ J/K} \cdot (290\text{K} + 627\text{K}) = 2.531 \cdot 10^{-15} \text{ W}$$

$$P_S = P_N \cdot \left(\frac{S}{N}\right) = 2.531 \cdot 10^{-15} \text{ W} \cdot 10 = 2.531 \cdot 10^{-14} \text{ W}$$

$$P_{S\text{dBm}} = 10 \log_{10} \frac{P_S}{1\text{mW}} = -106\text{dBm}$$

Poenostavljen izračun izključno v primeru $T_A \approx T_0 = 290\text{K}$

$$P_{S\text{dBm}} \approx (S/N)_{\text{dB}} + (\Delta f)_{\text{dB}\cdot\text{Hz}} + (k_B T_0)_{\text{dBm/Hz}} + F_{\text{dB}}$$

$$(k_B T_0)_{\text{dBm/Hz}} = 10 \log_{10} \frac{k_B T_0}{1\text{mJ}} \approx -174 \text{ dBm/Hz} \quad (\Delta f)_{\text{dB}\cdot\text{Hz}} = 10 \log_{10} \left(\frac{\Delta f}{1\text{Hz}}\right) = 53 \text{ dB}\cdot\text{Hz}$$

$$P_{S\text{dBm}} \approx 10\text{dB} + 53\text{dB}\cdot\text{Hz} - 174\text{dBm/Hz} + 5\text{dB} = -106\text{dBm}$$

Dva različna sprejemnika #1 in #2:

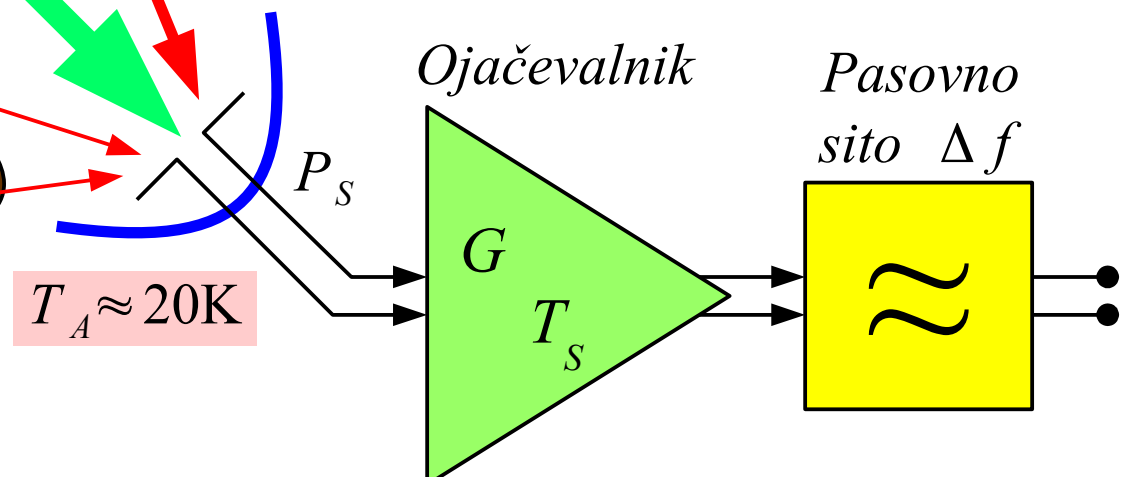
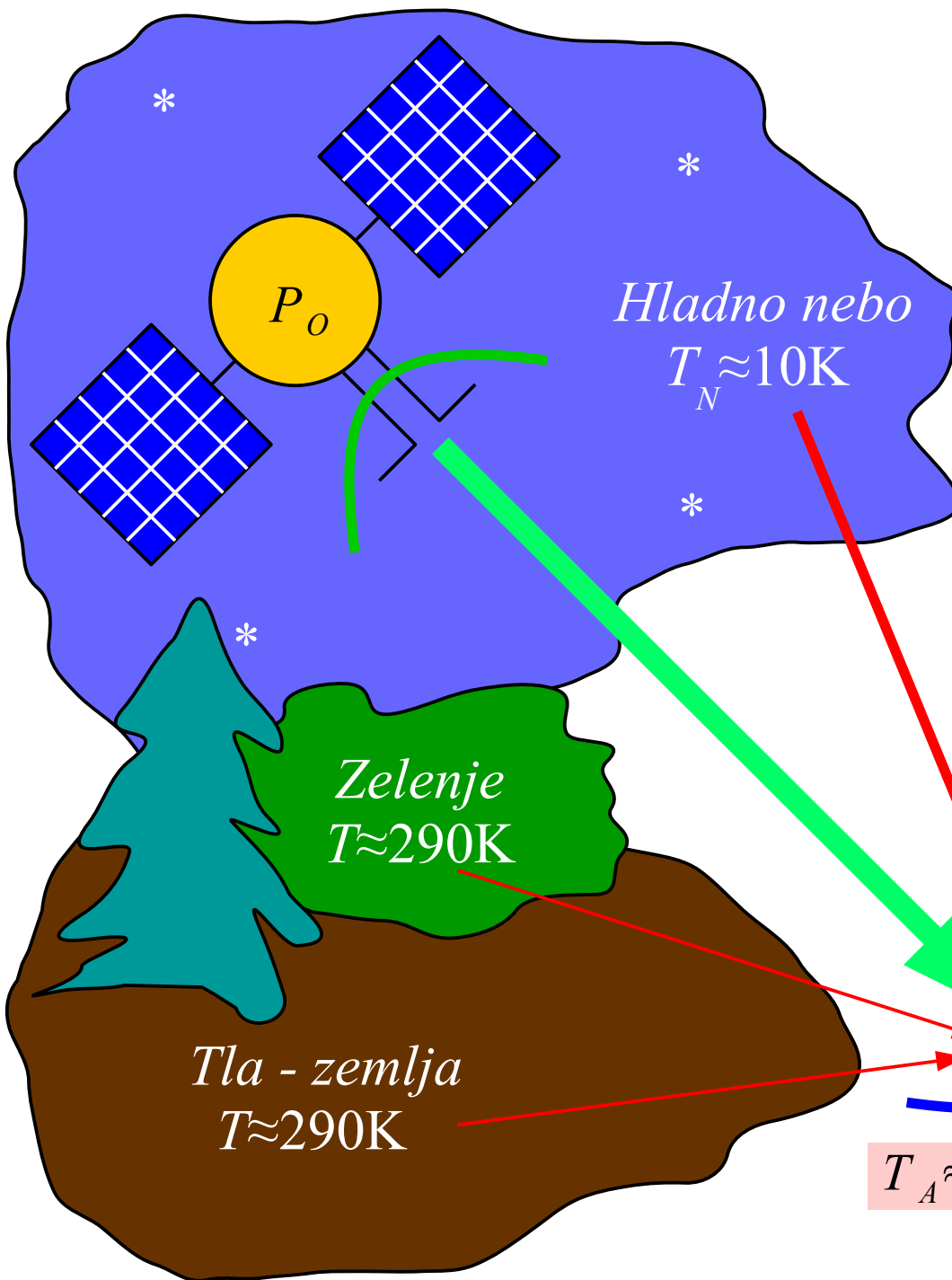
$$F_1 = 1\text{dB} \rightarrow T_1 = 75\text{K}$$

$$F_2 = 0.5\text{dB} \rightarrow T_2 = 35\text{K}$$

$$\Delta F_{\text{dB}} = F_1 - F_2 = 0.5\text{dB}$$

$$\Delta \left(\frac{S}{N} \right)_{\text{dB}} = 10 \log_{10} \left[\frac{T_A + T_2}{T_A + T_1} \right]$$

$$\Delta \left(\frac{S}{N} \right)_{\text{dB}} = 10 \log_{10} \left[\frac{20\text{K} + 75\text{K}}{20\text{K} + 35\text{K}} \right] = 2.37\text{dB}$$

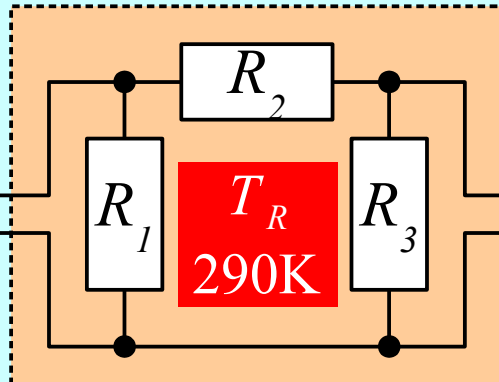


Signal generator

$P_o \approx 10\text{mW}$



Oscillator



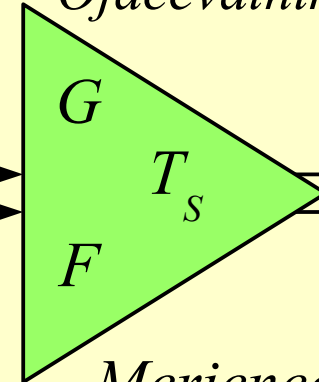
Slabilec a_{dB}

Oklop $> 150\text{dB}$!

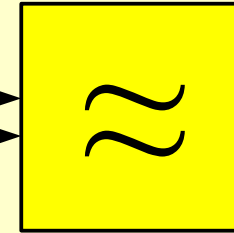
P_s

T_A

Ojačevalnik



Pasovno
sito Δf



Merjenec - sprejemnik

$\left(\frac{S}{N}\right)_{\text{izhod}}$

$50\text{dB} < a_{\text{dB}} < 150\text{dB}$

Sklop preko sevanja?

Dodatni zahtevi za merilni izvor (signal generator) za merjenje občutljivosti radijskih/mikrovalovnih sprejemnikov:

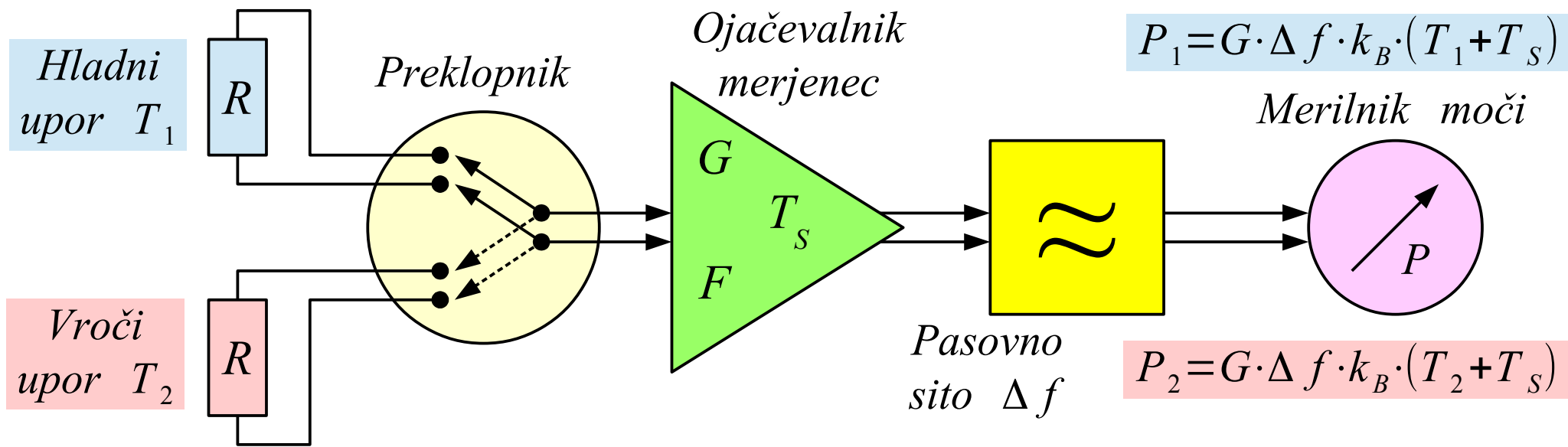
(1) Oklop $> 150\text{dB}$

(2) $T_R = T_A = T_0 = 290\text{K}$

(1) Zahtevani S/N pred demodulatorjem?

(2) Zahtevani S/N za demodulatorjem?

(3) Zahtevani BER?



V razmerju Y se neznanke $G \cdot \Delta f \cdot k_B$ natančno krajšajo!

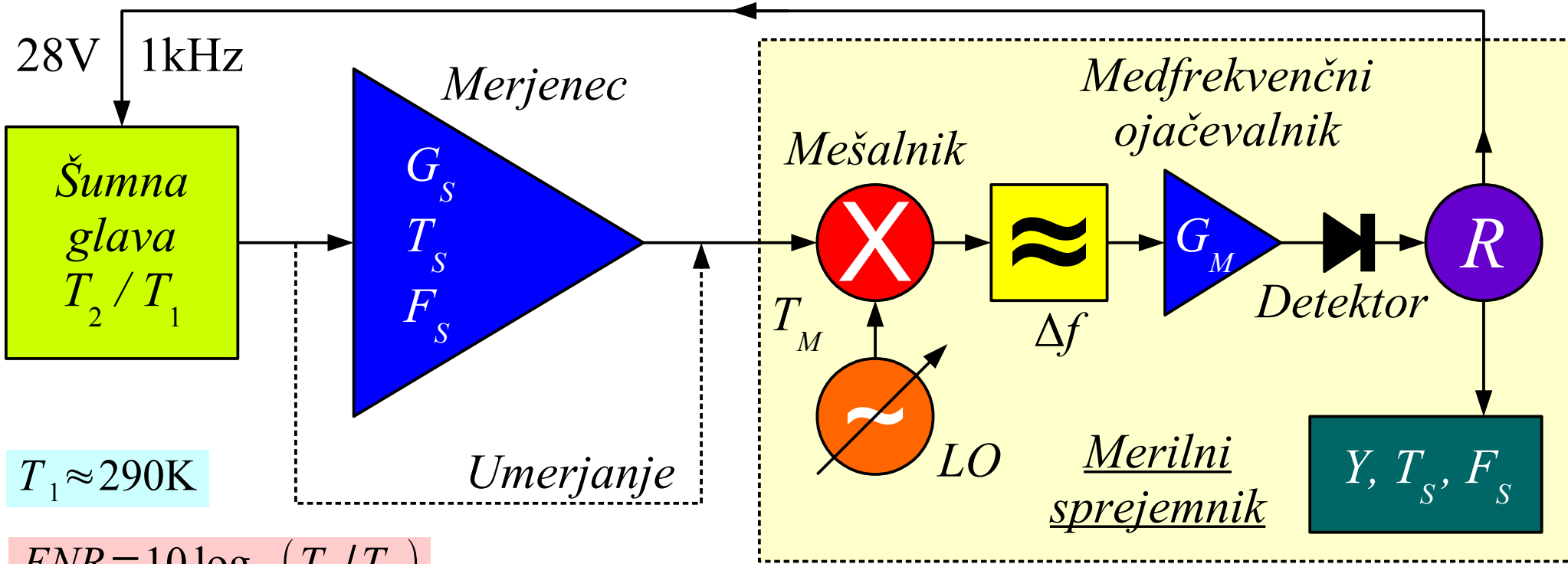
$$Y = \frac{P_2}{P_1} = \frac{T_2 + T_S}{T_1 + T_S}$$

$$T_S = \frac{T_2 - Y \cdot T_1}{Y - 1}$$

$$T_0 = 290\text{K}$$

$$(F_S)_{dB} = 10 \log_{10} \left[1 + \frac{T_2 - Y \cdot T_1}{(Y - 1) \cdot T_0} \right]$$

Vrsta upora	Temperatura
Antena v hladno nebo	$\sim 20\text{K}$
R hlajen tekoči dušik	$\sim 77\text{K}$
Antena v absorber	$\sim 290\text{K}$
R sobna temperatura	$\sim 290\text{K}$
Nitka žarnice kot R	$\sim 2000\text{K}$
Ioniziran plin kot R	$\sim 10^4\text{K}$
Plazovni preboj v diodi	$\sim 10^6\text{K}$



$T_1 \approx 290K$

$ENR = 10 \log_{10}(T_2/T_1)$

Dve meritvi brez umerjanja:

$$Y = \frac{P_2}{P_1} = \frac{T_2 + T_S + T_M / G_S}{T_1 + T_S + T_M / G_S}$$

$$T_S = \frac{T_2 - Y \cdot T_1}{Y - 1} - \frac{T_M}{G_S} \leftarrow \text{poznam } G_S$$

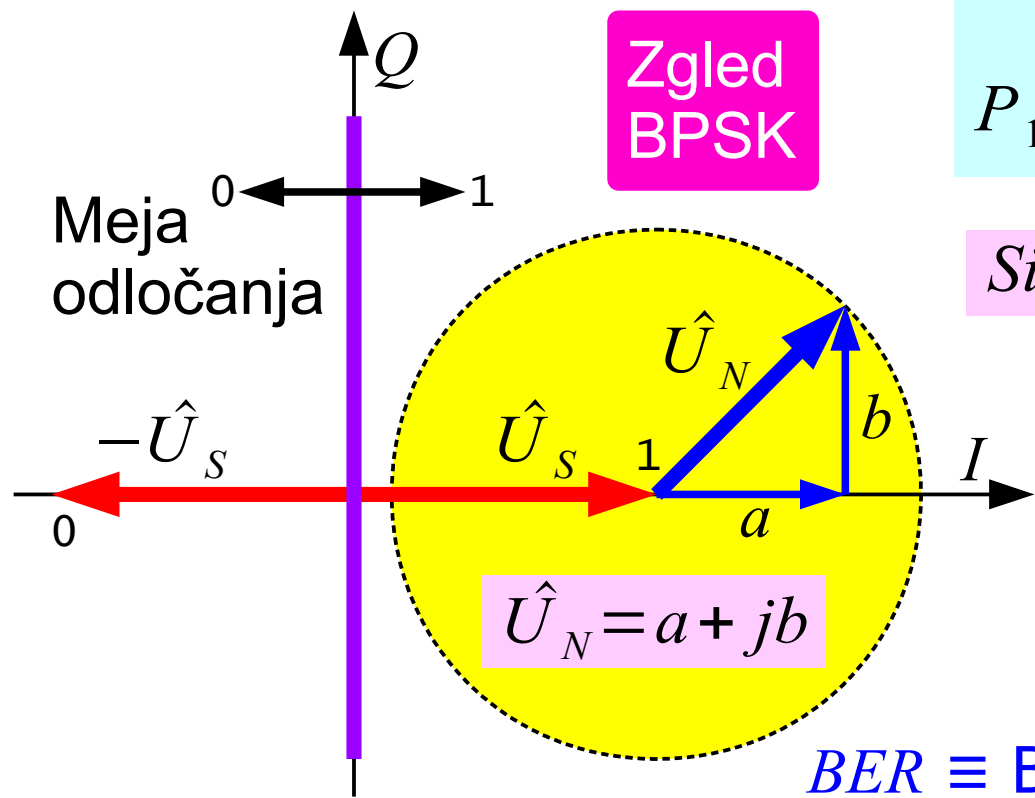
$$(F_S)_{dB} = 10 \log_{10} \left[1 + \frac{1}{T_0} \cdot \left(\frac{T_2 - Y \cdot T_1}{Y - 1} - \frac{T_M}{G_S} \right) \right]$$

$G_M \Delta f \equiv \text{nezanesljiv!}$

Štiri meritve z umerjanjem:

- (1) $P_1 = G_M G_S \Delta f k_B (T_1 + T_S + T_M / G_S)$
- (2) $P_2 = G_M G_S \Delta f k_B (T_2 + T_S + T_M / G_S)$
- (3) $P_3 = G_M \Delta f k_B (T_1 + T_M)$
- (4) $P_4 = G_M \Delta f k_B (T_2 + T_M)$

Rešujem 4 enačbe za 4 neznanke:
 T_S, G_S, T_M in $(G_M \Delta f k_B)$



Zgled BPSK

$$P_{1 \rightarrow 0} = \int_{-\infty}^{-|\hat{U}_s|} p(a) da$$

$$P_{0 \rightarrow 1} = \int_{|\hat{U}_s|}^{\infty} p(a) da$$

Simetrična meja: $P_{1 \rightarrow 0} = P_{0 \rightarrow 1} = BER$

$$BER = \int_{|\hat{U}_s|}^{\infty} \frac{1}{\sqrt{\pi} \langle |\hat{U}_N|^2 \rangle} e^{-\frac{a^2}{\langle |\hat{U}_N|^2 \rangle}} da$$

$$\text{erfc}(x) = \frac{2}{\sqrt{\pi}} \int_x^{\infty} e^{-u^2} du$$

Gaussova porazdelitev gostote verjetnosti sofazne a in kvadraturene b komponente šuma

$$p(a) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{a^2}{2\sigma^2}}$$

$$p(b) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{b^2}{2\sigma^2}}$$

$$BER = \frac{1}{2} \text{erfc} \left(\frac{|\hat{U}_s|}{\sqrt{\langle |\hat{U}_N|^2 \rangle}} \right)$$

$$P_S = \alpha |\hat{U}_s|^2$$

$$P_N = \alpha \langle |\hat{U}_N|^2 \rangle$$

$$\langle |\hat{U}_N|^2 \rangle = \langle a^2 \rangle + \langle b^2 \rangle = 2\sigma^2$$

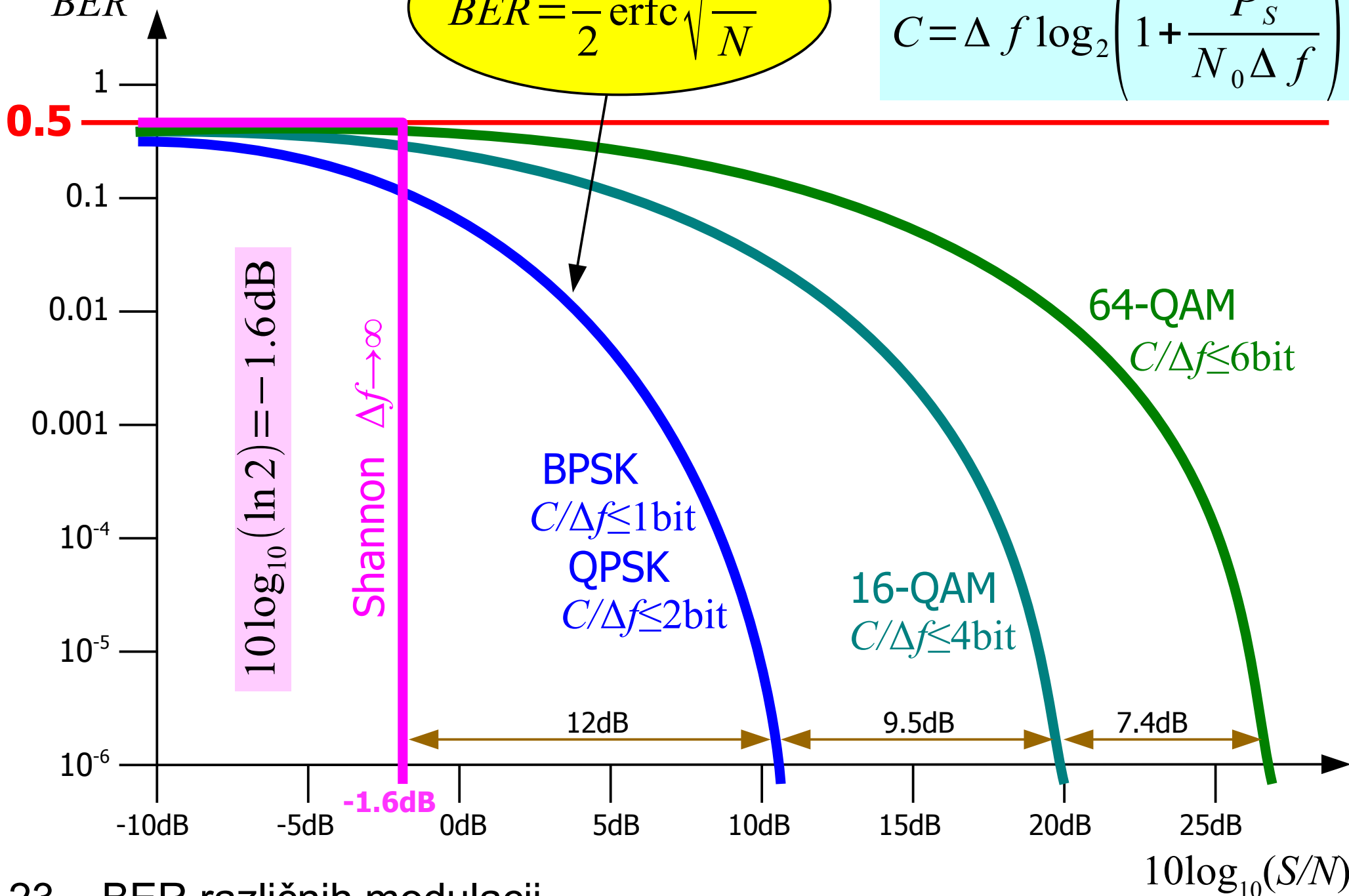
$$BER = \frac{1}{2} \text{erfc} \left(\sqrt{\frac{P_S}{P_N}} \right)$$

Pogostnost napak
BER

$$BER = \frac{1}{2} \operatorname{erfc} \sqrt{\frac{S}{N}}$$

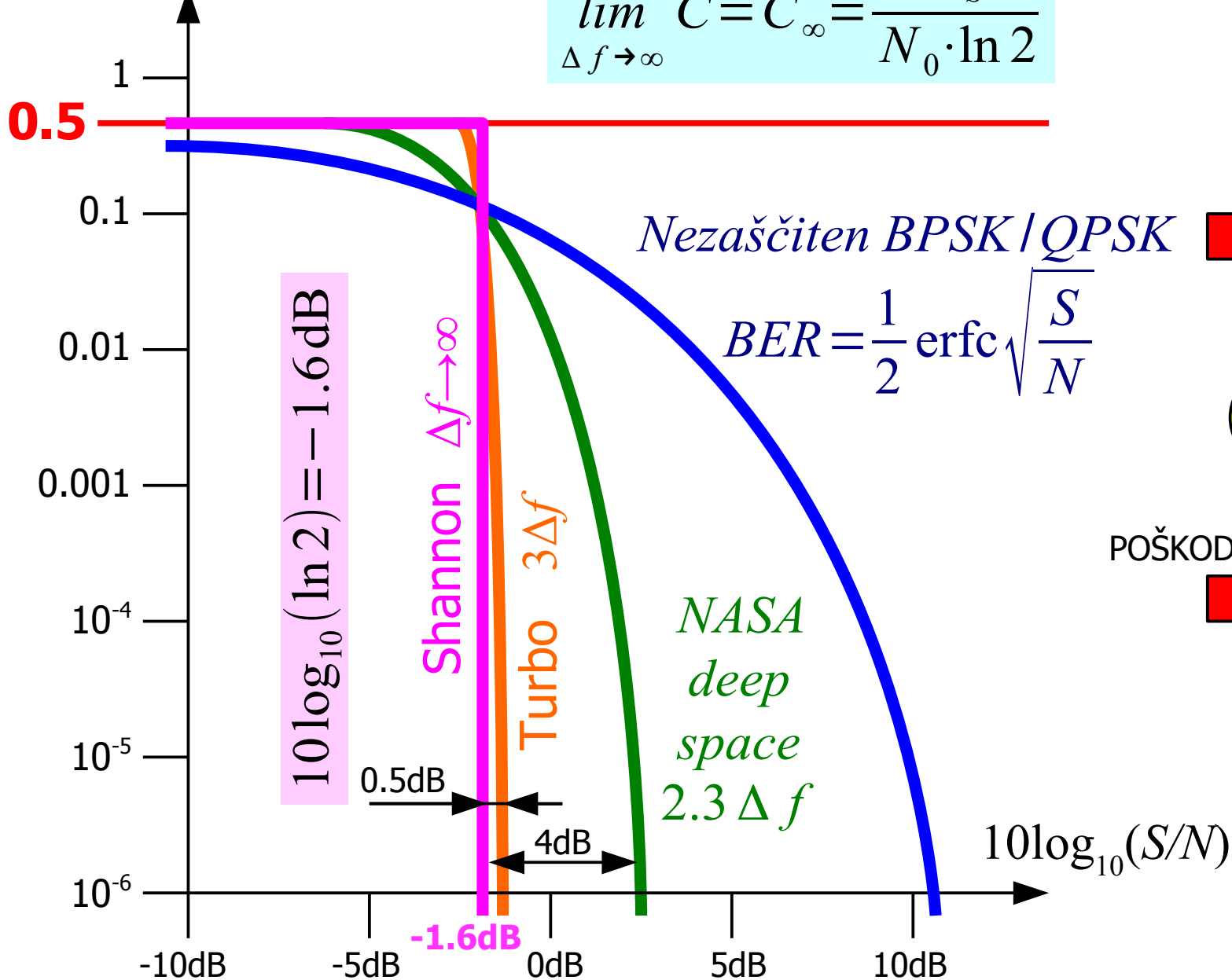
Shannon

$$C = \Delta f \log_2 \left(1 + \frac{P_s}{N_0 \Delta f} \right)$$



Pogostnost napak

BER



Shannon

$$\lim_{\Delta f \rightarrow \infty} C = C_{\infty} = \frac{P_s}{N_0 \cdot \ln 2}$$

Nezaščiten BPSK/QPSK

$$BER = \frac{1}{2} \text{erfc} \sqrt{\frac{S}{N}}$$

*NASA
deep
space
2.3 Δf*

SPOROČILO

FEC
KODER

SPOROČILO

PARITETA

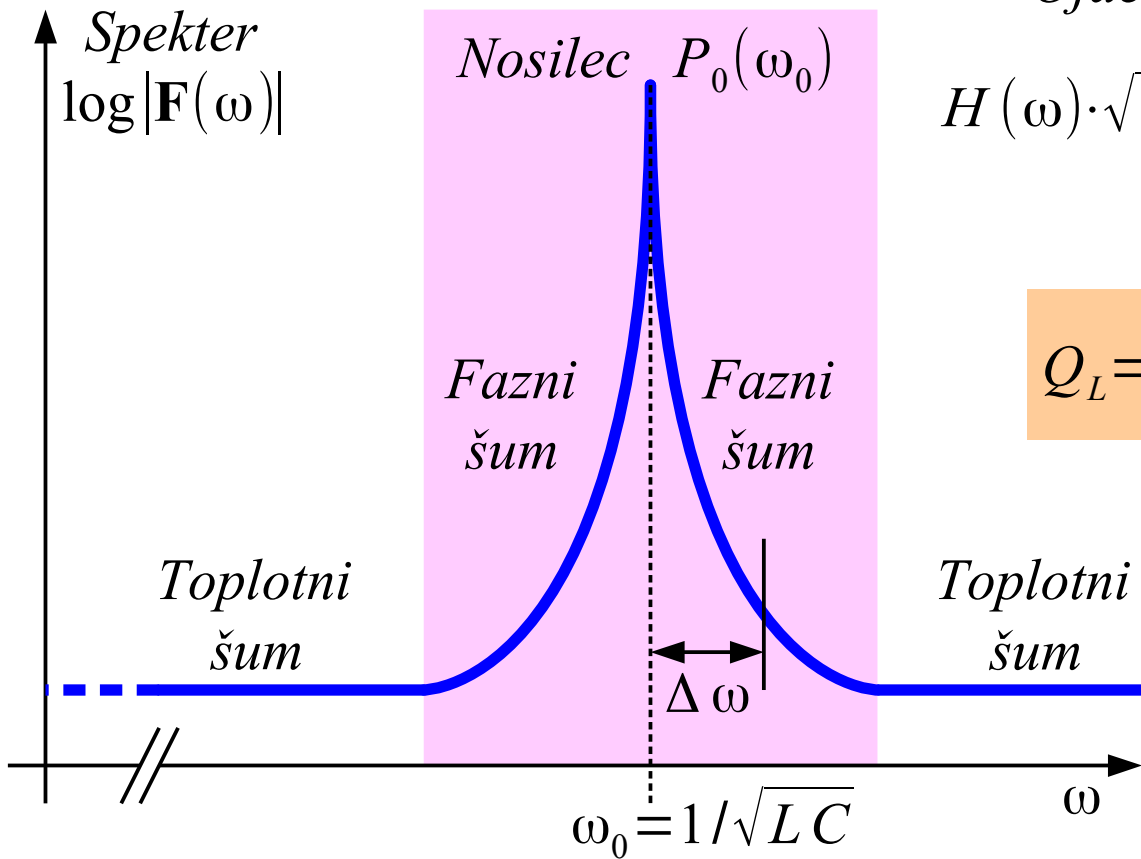
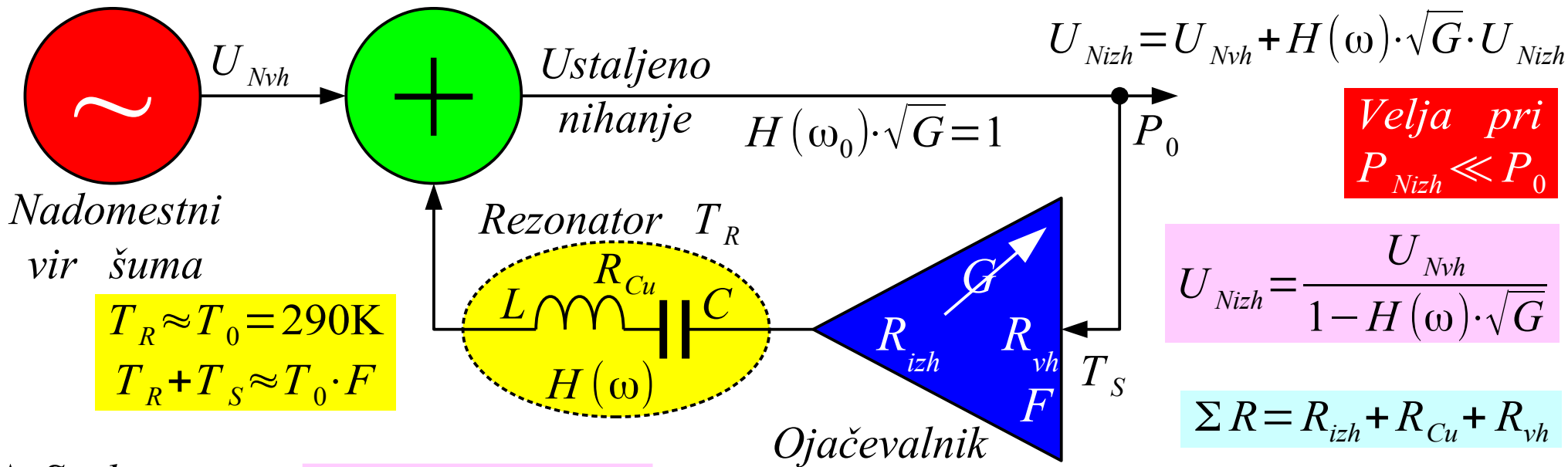
IZGUBNA
PRENOSNA
POT

POŠKODOVANO!

FEC
DEKODER

SPOROČILO

POPRAVLJENO!



$$H(\omega) \cdot \sqrt{G} = \frac{\Sigma R}{\Sigma R + j\omega L + \frac{1}{j\omega C}} \approx \frac{1}{1 + j2Q_L \frac{\Delta\omega}{\omega_0}}$$

$$Q_L = \frac{\omega_0 L}{\Sigma R}$$

$$U_{Nizh} \approx U_{Nvh} \cdot \left(1 + \frac{\omega_0}{j2Q_L \Delta\omega} \right)$$

$$P_{Nizh} \approx P_{Nvh} \cdot \left[1 + \left(\frac{\omega_0}{2Q_L \Delta\omega} \right)^2 \right]$$

Amplitudni in fazni šum

$$P_{Nizh} \approx P_{Nvh} \cdot \left[1 + \left(\frac{f_0}{2Q_L \Delta f} \right)^2 \right]$$

Normirana
spektralna gostota
faznega šuma

$P_0 \equiv$ moč nosilca

$$\frac{dP_{Nvh}}{df} = N_0 = k_B(T_R + T_S) \approx k_B T_0 F$$

$\log L(\Delta f)$
[dBc/Hz]

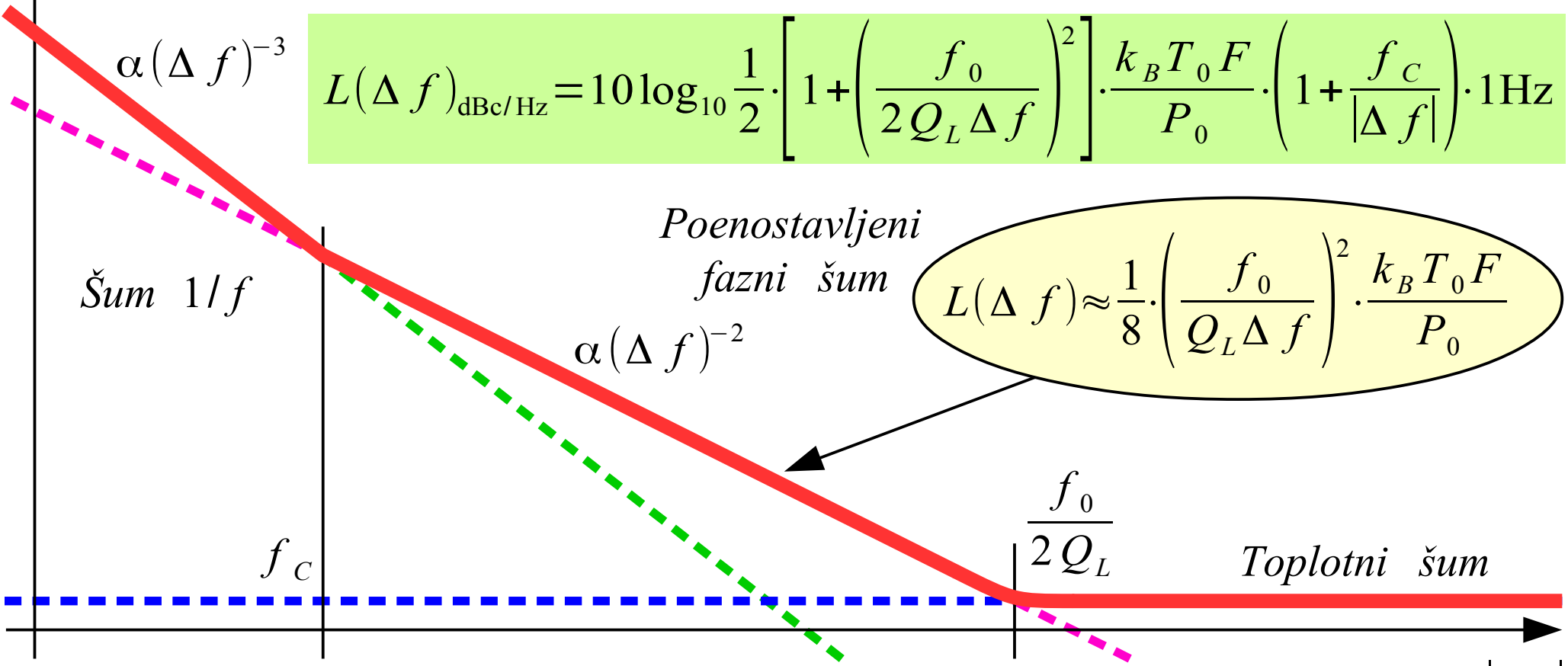
$$L(\Delta f) = \frac{1}{P_0} \cdot \frac{dP_\varphi}{df} = \frac{1}{2} \cdot \left[1 + \left(\frac{f_0}{2Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left(1 + \frac{f_c}{|\Delta f|} \right) \quad [\text{Hz}^{-1}]$$

Velja pri
 $L(\Delta f) \cdot \Delta f \ll 1$

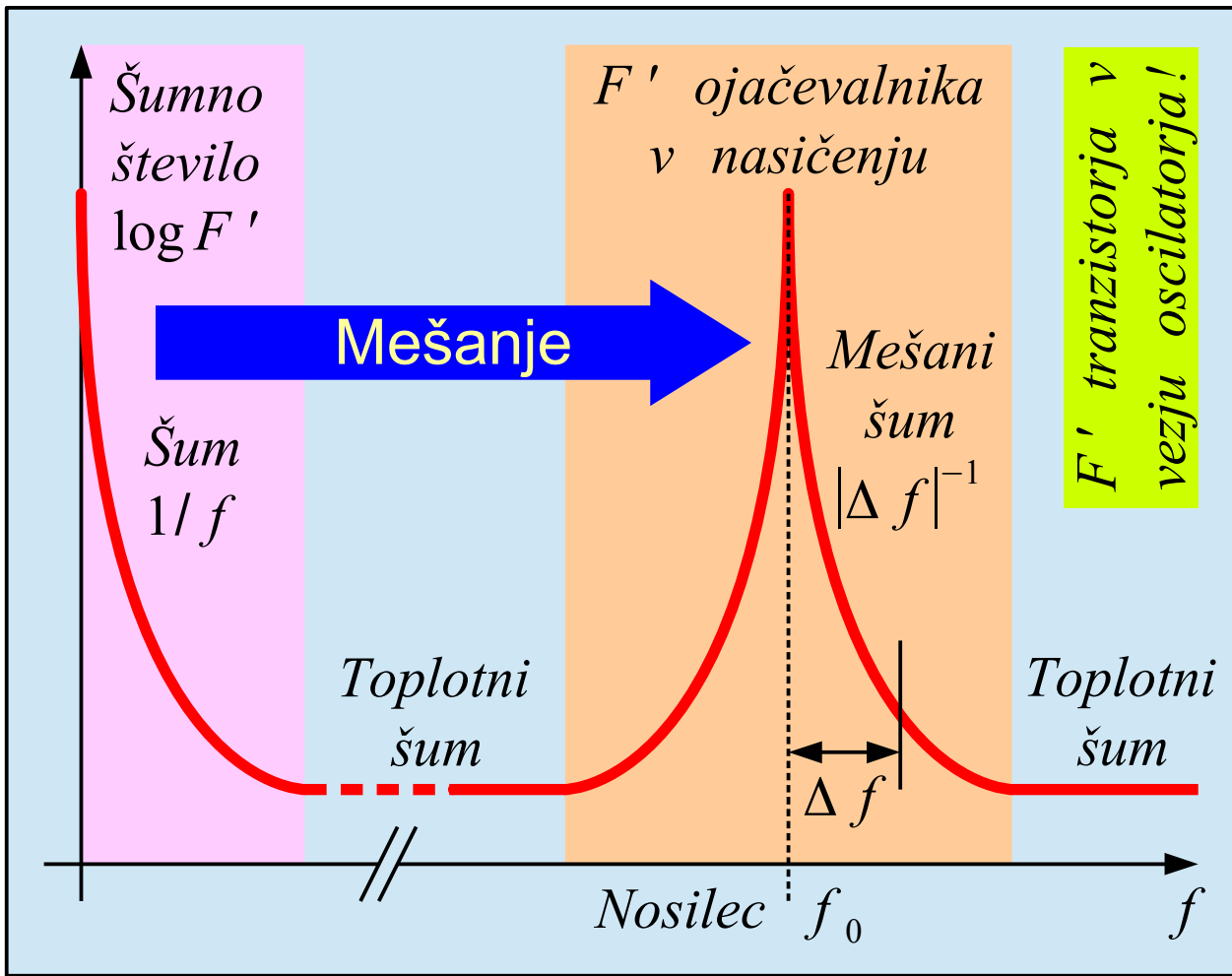
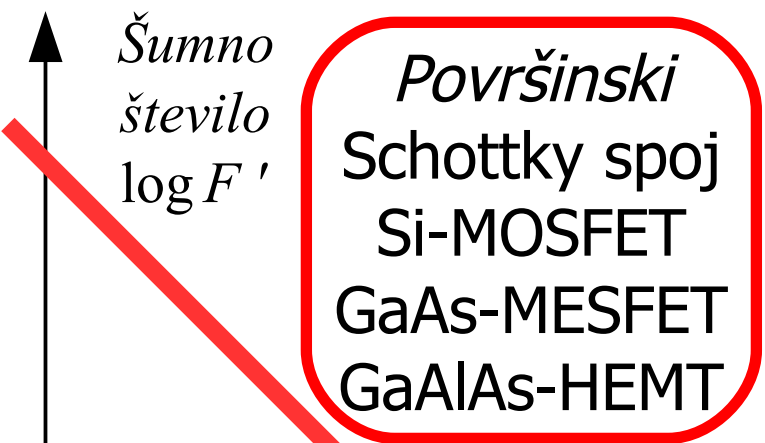
Samo fazni šum

Nasičenje odstrani
amplitudni šum

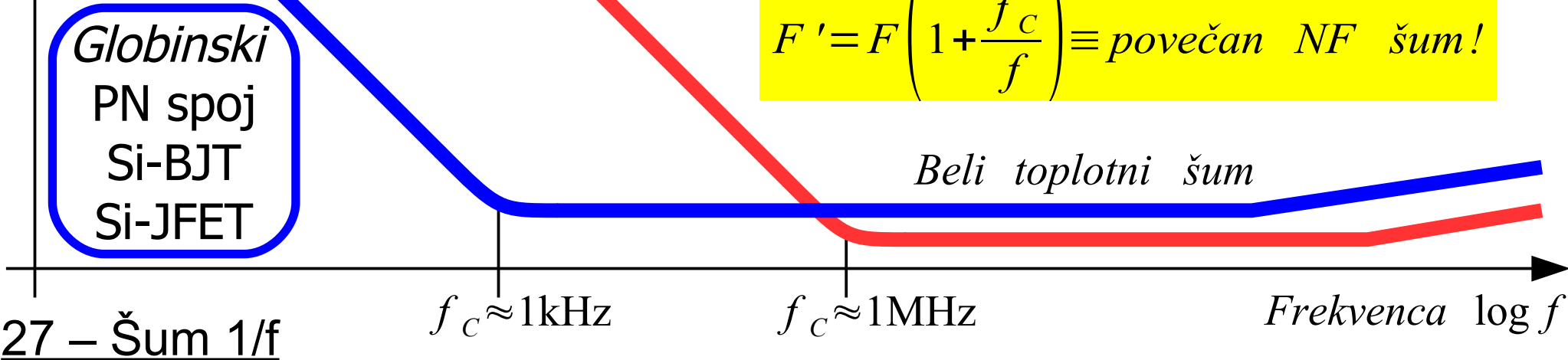
Šum $1/f$



Šum $1/f$ običajno nima jasne fizikalne razlage!



$$F' = F \left(1 + \frac{f_c}{f} \right) \equiv \text{povečan NF šum!}$$

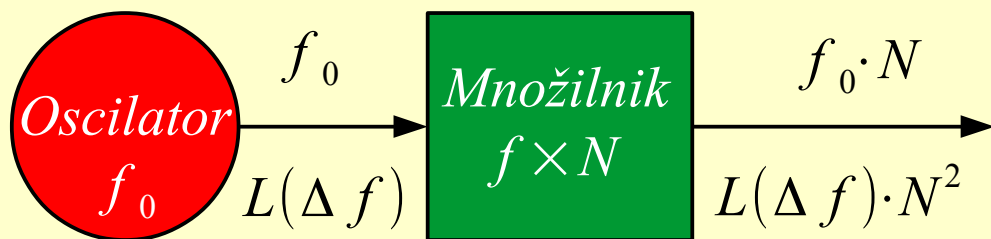


Kvaliteta obremenjenega rezonatorja Q_L je ključnega pomena za fazni šum!

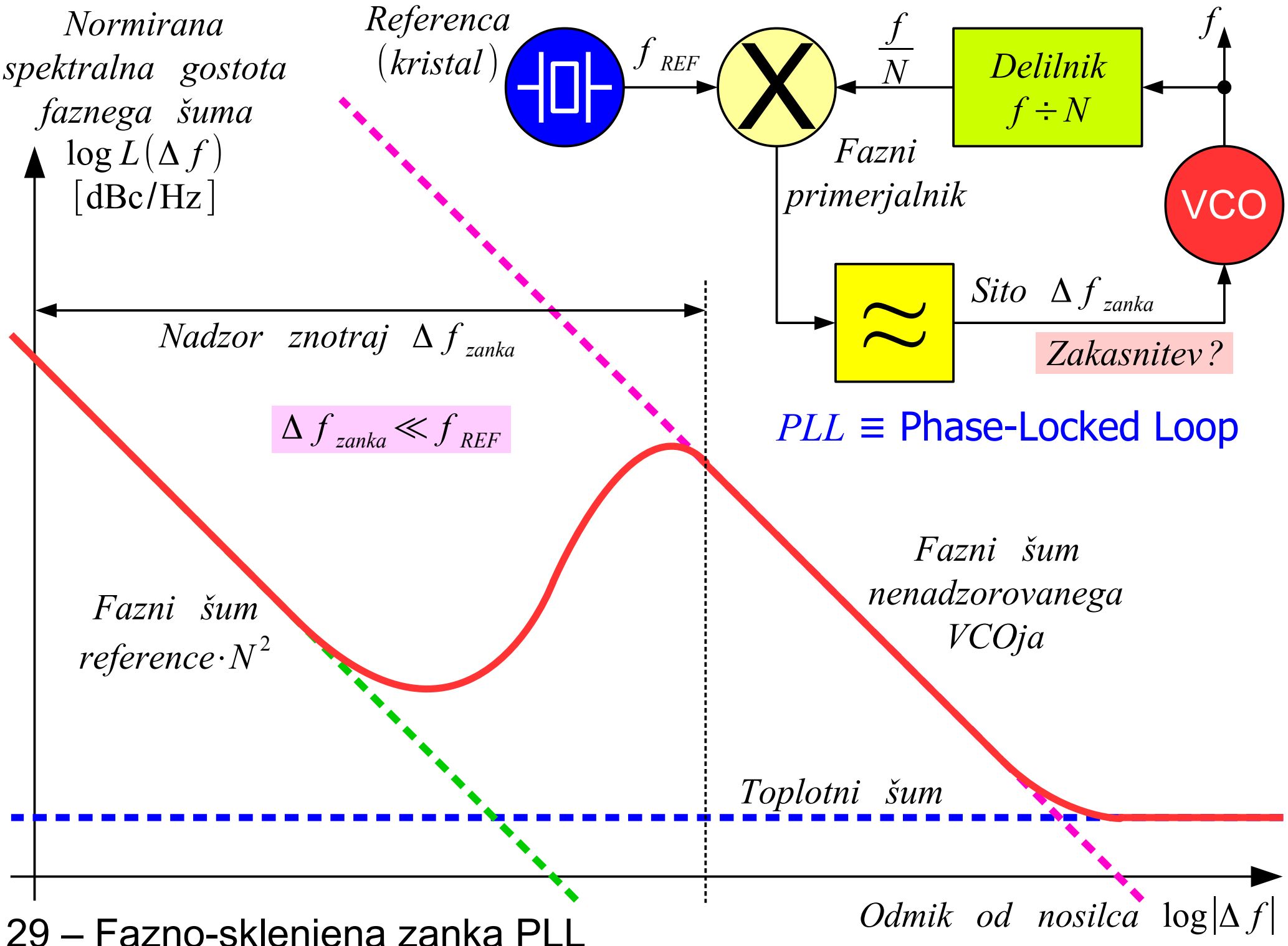
$$L(\Delta f) = \frac{1}{2} \cdot \left[1 + \left(\frac{f_0}{2 Q_L \Delta f} \right)^2 \right] \cdot \frac{k_B T_0 F}{P_0} \cdot \left(1 + \frac{f_c}{|\Delta f|} \right)$$

Frekvenčno nastavljivi oscilatorji	Q_L
RC VCO	~ 1
Cev BWO	~ 1
Varikap LC VCO	$10 \leftrightarrow 30$
YIG ($Y_3Fe_5O_{12}$) oscilator	$300 \leftrightarrow 1000$

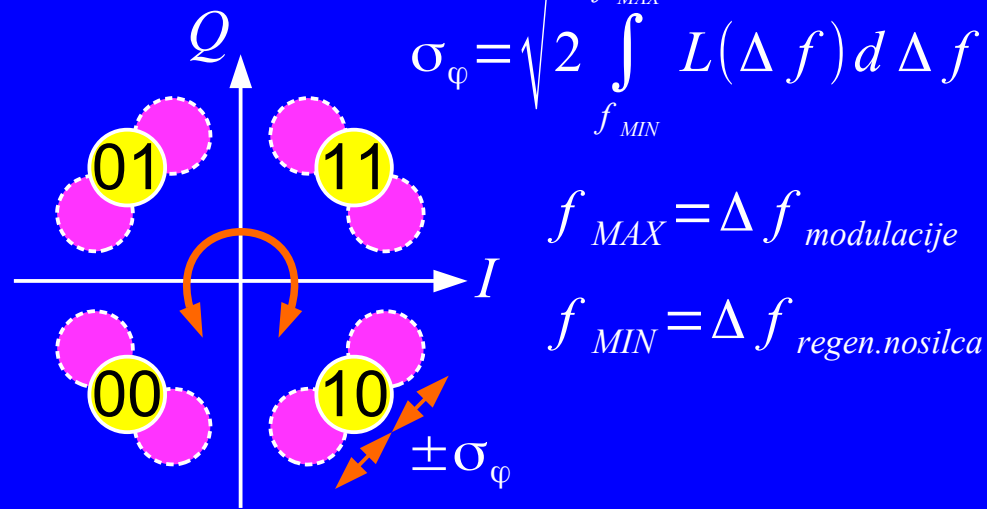
Oscilatorji fiksne frekvence	Q_L
RC multivibrator	~ 1
LC nihajni krog	$30 \leftrightarrow 100$
Votlinski rezonator	$1000 \leftrightarrow 3000$
Keramični dielektrični rezonator	$1000 \leftrightarrow 3000$
AT kremenov kristal (osnovna rezonanca)	$3000 \leftrightarrow 10000$
AT kremenov kristal (tretji/peti overton)	$10000 \leftrightarrow 30000$
Elektro-optični zakasnilni vod (\$)	$\sim 10^5$
Safirjev dielektrični rezonator (\$\$\$)	$\sim 3 \cdot 10^5$
Rdeč HeNe LASER	$\sim 10^8$



Fazni šum se množi s kvadratom množenja frekvence!

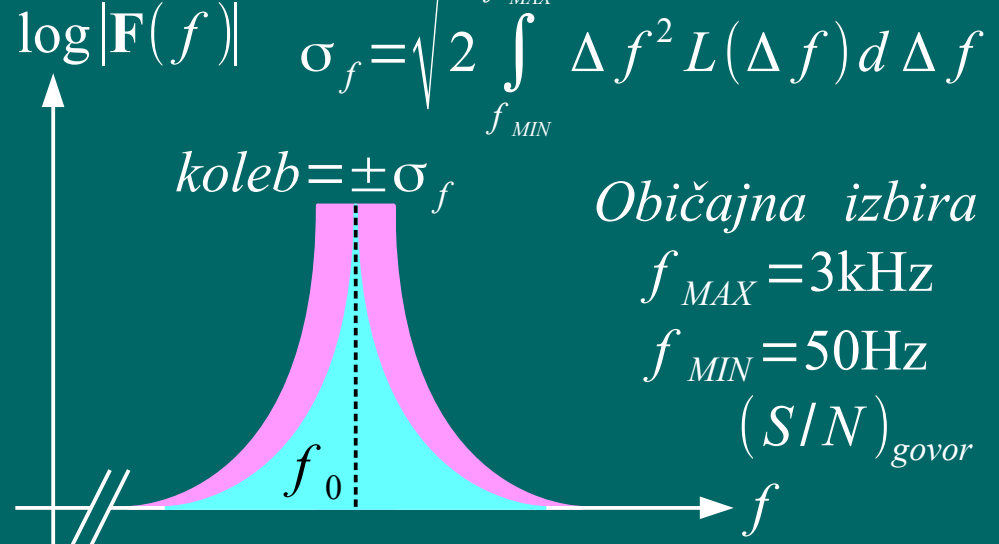


Zgled QPSK



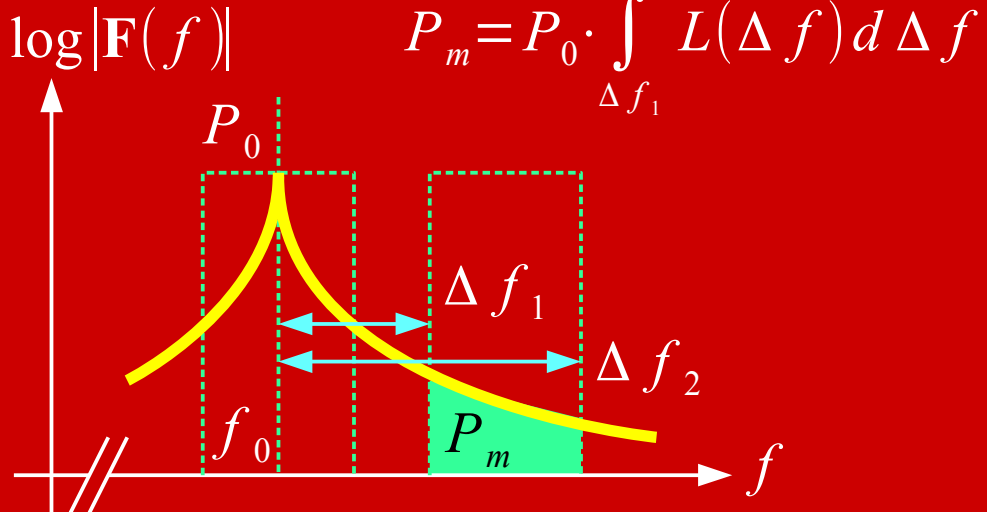
Zasuk ozvezdja modulacije

Spekter

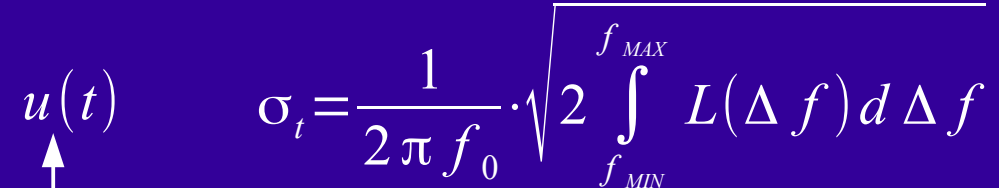


Naključna FM (residual FM)

Spekter

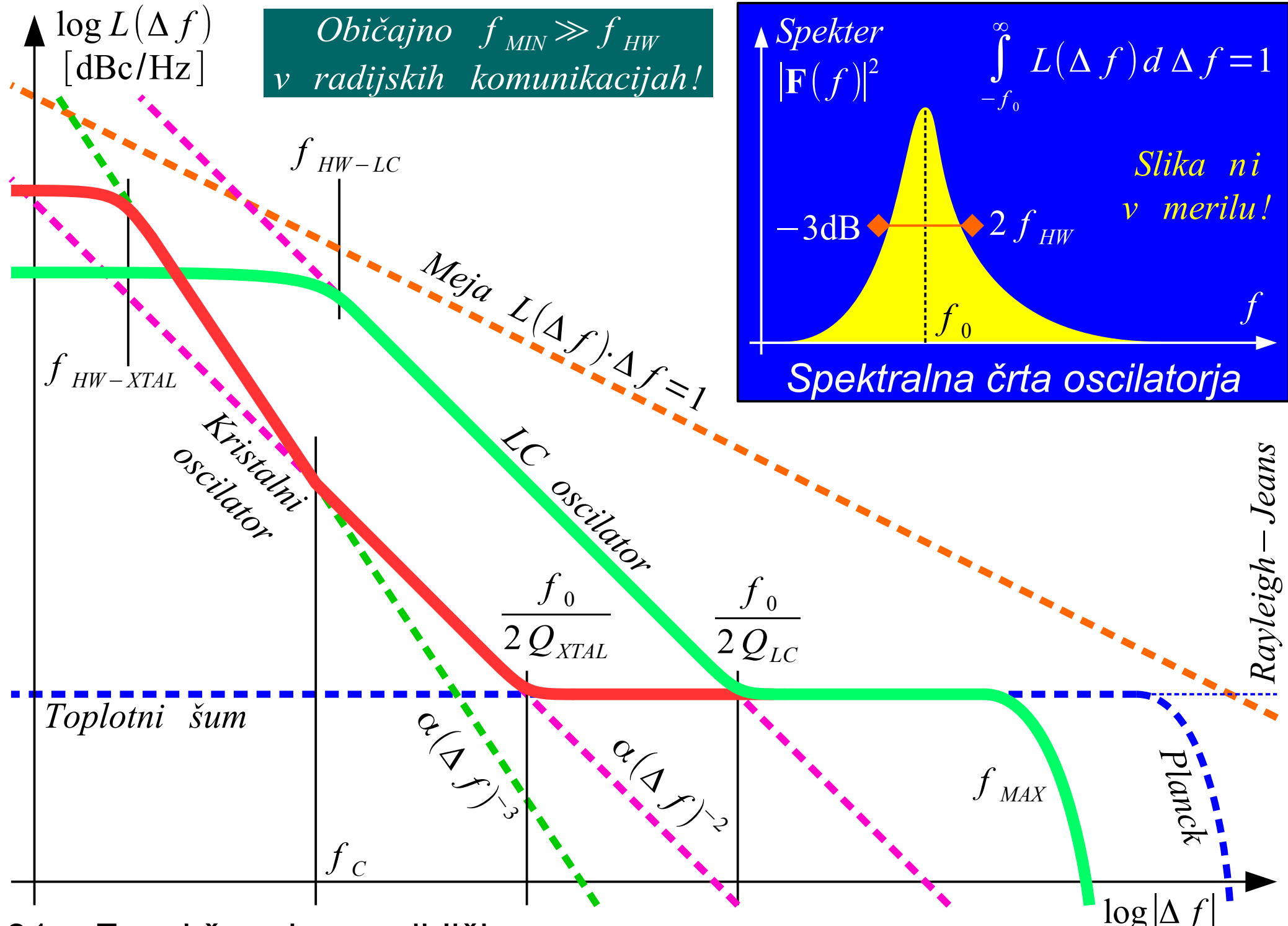


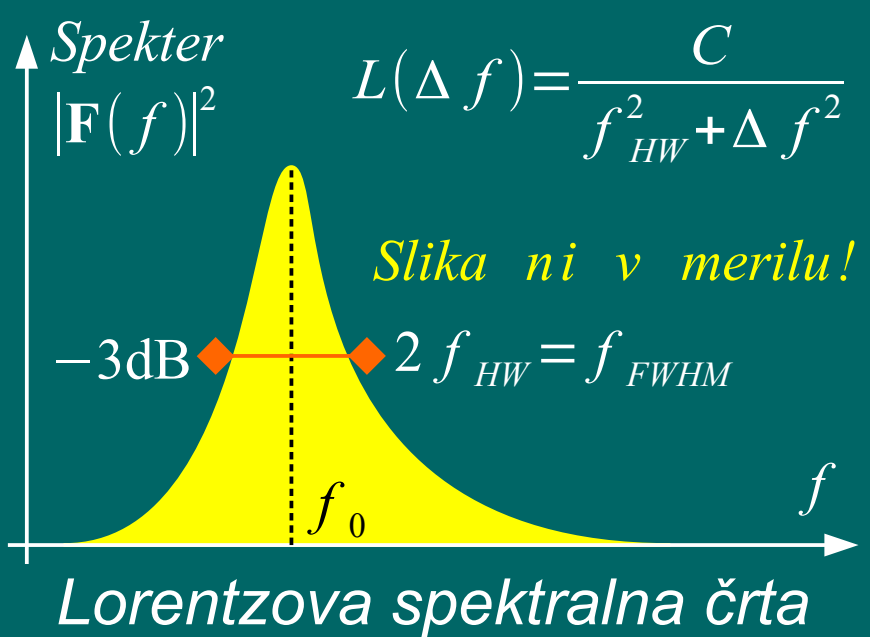
Motnja v sosednjem kanalu



$f_{MAX} = \Delta f_{opazovanja}$ $f_{MIN} = 1/t_{opazovanja}$

Drhtenje ure (jitter)





*Prispevek toplotnega šuma je zanemarljiv
 f_{MAX} vezja oziroma Planckov zakon*

Šum $1/f$ LC oscilatorja je zanemarljiv

$$L(\Delta f) = \frac{1}{8} \cdot \left(\frac{f_0}{Q_L} \right)^2 \cdot \frac{1}{f_{HW}^2 + \Delta f^2} \cdot \frac{k_B T_0 F}{P_0}$$

Lorentzova črta v Leesonovi enačbi

$$\int_{-f_0}^{\infty} L(\Delta f) d\Delta f = 1 \approx \int_{-\infty}^{\infty} L(\Delta f) d\Delta f = \frac{1}{8} \cdot \left(\frac{f_0}{Q_L} \right)^2 \cdot \frac{k_B T_0 F}{P_0} \int_{-\infty}^{\infty} \frac{1}{f_{HW}^2 + \Delta f^2} d\Delta f =$$

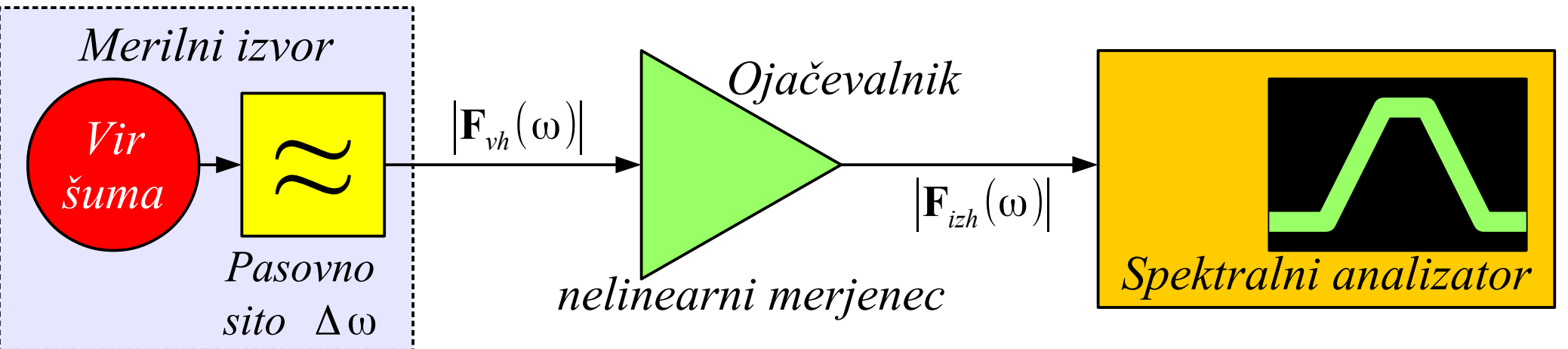
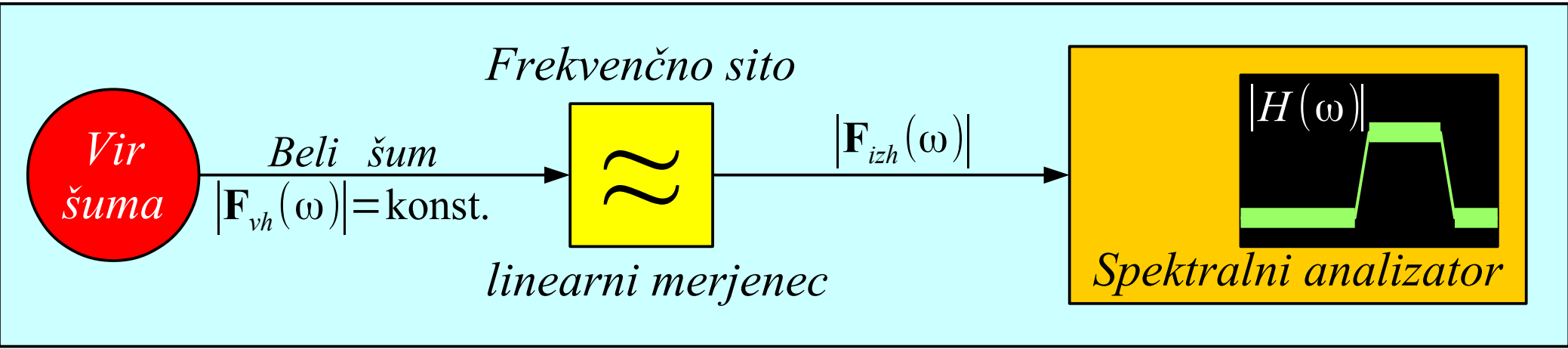
$$= \frac{1}{8} \cdot \left(\frac{f_0}{Q_L} \right)^2 \cdot \frac{k_B T_0 F}{P_0} \cdot \left[\frac{1}{f_{HW}} \cdot \arctan \frac{\Delta f}{f_{HW}} \right]_{\Delta f = -\infty}^{\Delta f = \infty} = \frac{k_B T_0 F}{8 P_0} \cdot \left(\frac{f_0}{Q_L} \right)^2 \cdot \frac{\pi}{f_{HW}}$$

$$f_{HW} = \frac{\pi k_B T_0 F}{8 P_0} \cdot \left(\frac{f_0}{Q_L} \right)^2$$

*Zgled $f_0 = 3\text{GHz}$ $Q_L = 10$
 $P_0 = 0.1\text{mW}$ $F = 10\text{dB}$
 $f_{HW} = 14\text{Hz}$ $f_{FWHM} = 28\text{Hz}$*

$$C = \frac{k_B T_0 F}{8 P_0} \cdot \left(\frac{f_0}{Q_L} \right)^2 = \frac{f_{HW}}{\pi}$$

$$L(\Delta f) = \frac{f_{HW} / \pi}{f_{HW}^2 + \Delta f^2}$$



Naravni vir naključnega signala:

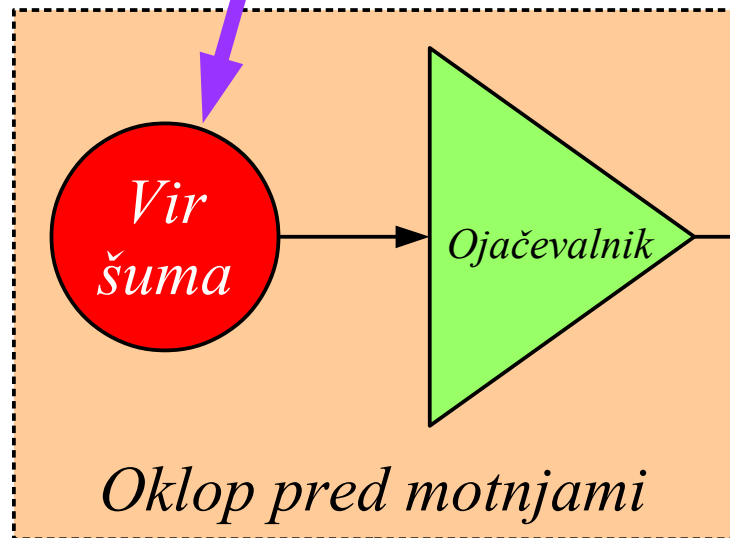
Toplotni šum

Zrnati šum

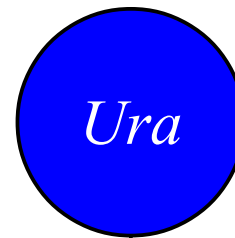
Plazovni preboj

Radioaktivni razpad

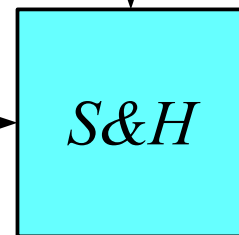
...



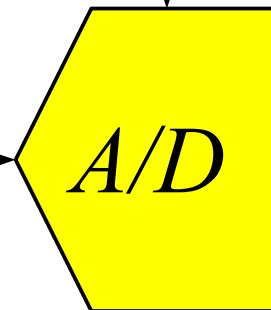
*Motnje niso naključne!
Motnje so lahko namerne!*



$$f_{ura} \ll \Delta f$$



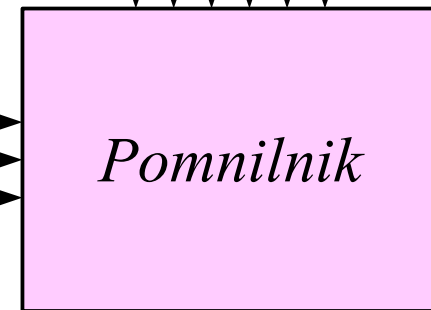
Vzorčenje



Kvantizacija



Števec naslovov



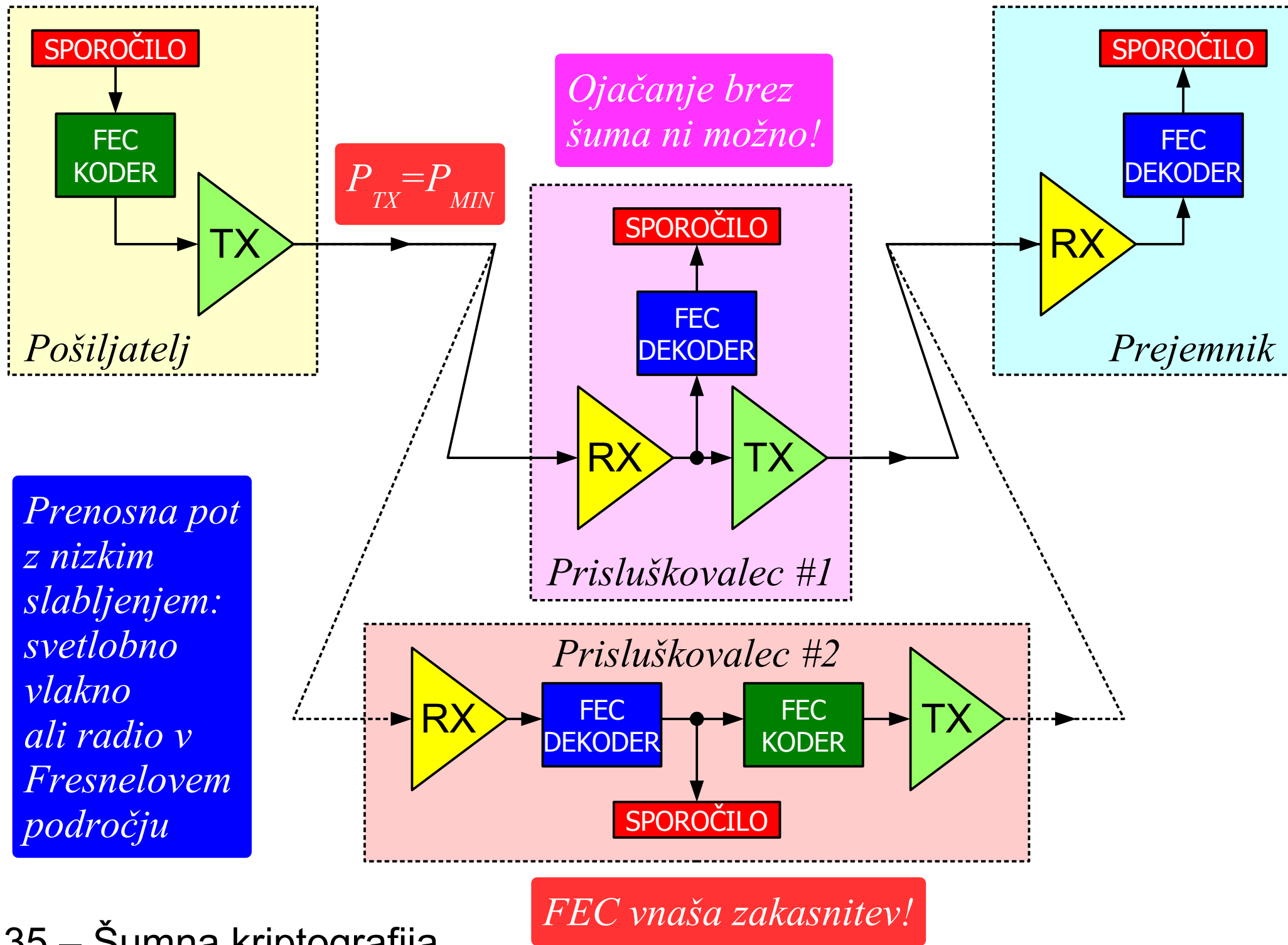
Pomnilnik

Kriptografski ključ poljube dolžine:

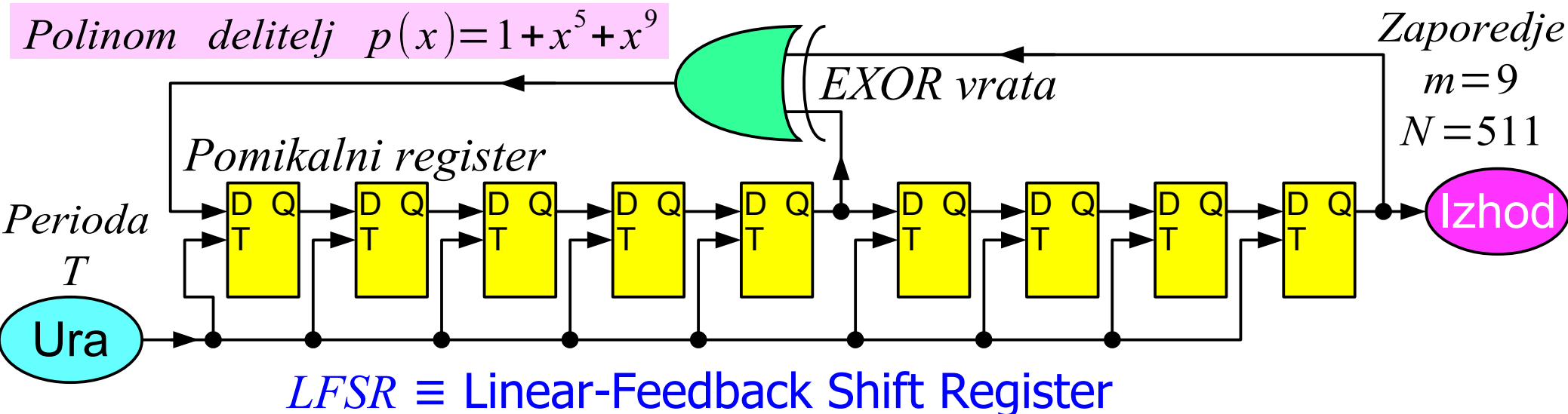
Geslo (razmeroma kratko)

Ključ DES, AES itd

Šifrirna knjiga za enkratno uporabo
(zelo dolga, a nezlomljiva šifra)...

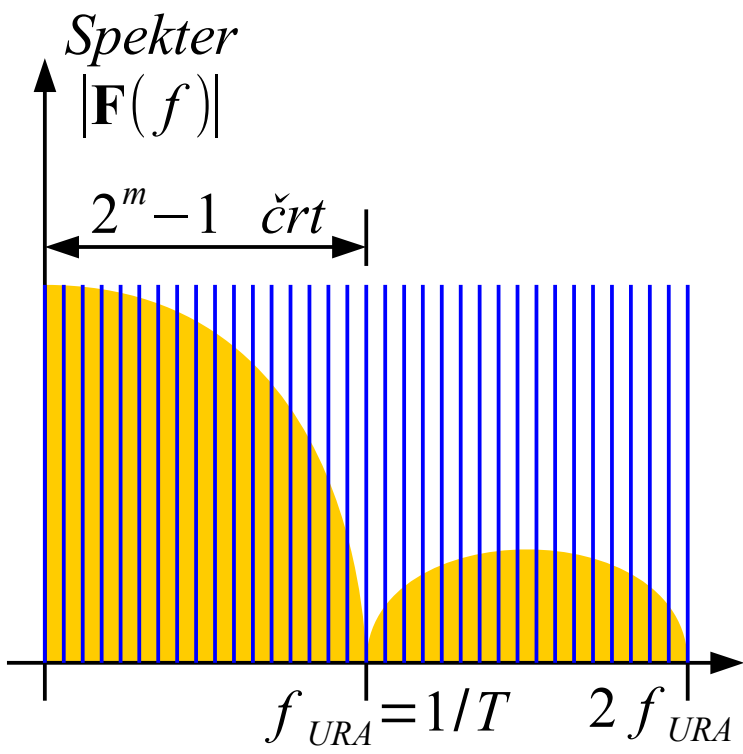
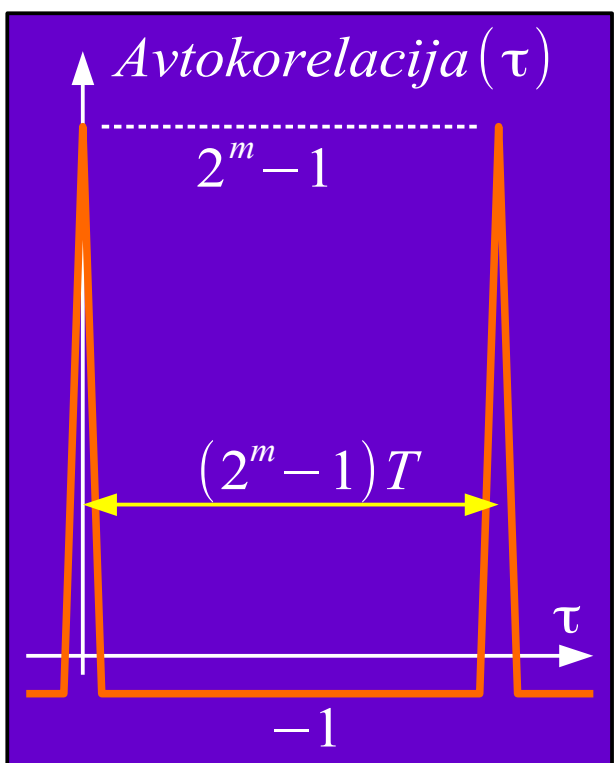


Polinom delitelj $p(x) = 1 + x^5 + x^9$



Nerazcepni polinom $p(x) = 1 + x^l + x^m \rightarrow$ zaporedje dolžine $\max N = 2^m - 1$

- 2^{m-1} enic in $2^{m-1} - 1$ ničel razporejenih v skupine
- 1X m enic, m-1 ničel
- 1X m-2 enic in ničel
- 2X m-3 enic in ničel
- 4X m-4 enic in ničel
-
- 2^{m-5} skupin 111 in 000
- 2^{m-4} skupin 11 in 00
- 2^{m-3} posamičnih 1 in 0



Sliši in vidi se kot beli šum!

Avtokorelacija ima dve vrednosti z enim samim izrazitim vrhom:

- zaporedja za sinhronizacijsko glavo podatkovnih okvirjev
- razširitvena zaporedja v CDMA
- natančen prenos časa v radionavigaciji (GPS, GLONASS)

Brezhiben spekter enako velikih, enakomerno razmaknjenih črt ter preprosto proizvodnje/preverjanje:

- preizkusni podatki za vse vrste zvez v telekomunikacijah
- skrambliranje (randomization) podatkov kot linijsko kodiranje

Razmerje vršna moč / povprečje:

$$LFSR: \frac{P_{MAX}}{\langle P \rangle} \approx 1 \quad \text{\textit{\textless}um: } \frac{P_{MAX}}{\langle P \rangle} \rightarrow \infty$$

Psevdonaključna zaporedja LFSR nimajo kriptološke vrednosti: algoritem Berlekamp-Massey 1969

Zaporedja LFSR so plod človeškega duha, najčistejša matematika, ki v naravi nikjer ne nastopa!

*Kako naj se prestavimo prebivalcem sosednje galaksije?
Kako ugotovimo, da nas oni iščejo?*